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2 **Nonlinear Data Assimilation by Deep Learning**

3 **Embedded in an Ensemble Kalman Filter**

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## Abstract

Recent progress in the particle filter has made it possible to use it for nonlinear or non-Gaussian data assimilation in high-dimensional systems, but a relatively large ensemble is still needed to outperform the ensemble Kalman filter (EnKF) in terms of accuracy. An alternative ensemble data assimilation method based on deep learning is presented, in which deep neural networks are locally embedded in the EnKF. This method is named the deep learning-ensemble Kalman filter (DL-EnKF). The DL-EnKF analysis ensemble is generated from the DL-EnKF analysis and the EnKF analysis deviation ensemble. The performance of the DL-EnKF is investigated through data assimilation experiments in both perfect and imperfect model scenarios using three versions of the Lorenz 96 model and a deterministic EnKF with an ensemble size of 10. Nonlinearity in data assimilation is controlled by changing the time interval between observations. Results demonstrate that despite such a small ensemble the DL-EnKF is superior to the EnKF in terms of accuracy in strongly nonlinear regimes and that the DL-EnKF analysis is more accurate than the output of deep learning due to positive feedback in assimilation cycles. Even if the target of training is an EnKF analysis with a large ensemble or a simulation by an imperfect model, the improvement introduced by the DL-EnKF is not very different from the case where the target of training is the true state.

51 **Keywords** data assimilation; deep learning; ensemble Kalman filter; particle filter

52

## 53 **1. Introduction**

54 Data assimilation in nonlinear or non-Gaussian systems has been a challenge in  
55 meteorology and other geosciences (Bocquet et al., 2010). For instance, it is well known  
56 that cumulus convection exhibits strong non-Gaussianity in data assimilation (e.g., Kondo  
57 and Miyoshi, 2019; Kawabata and Ueno, 2020). The ensemble Kalman filter (EnKF) is  
58 formulated under the Gaussian assumption and is close to optimal in weakly nonlinear  
59 regimes, but it does not work well if nonlinearity is strong. On the other hand, the particle  
60 filter (PF) does not need the Gaussian assumption, but the weight degeneracy had been  
61 preventing the use of the PF for high-dimensional data assimilation (Snyder et al., 2008; van  
62 Leeuwen 2009). However, this limitation is disappearing due to recent developments in the  
63 PF, including the use of localization and hybrids with the EnKF (Farchi and Bosquet, 2018;  
64 van Leeuwen et al., 2019). Despite this progress, a relatively large ensemble is still needed  
65 for the PF to outperform the EnKF (e.g., Penny and Miyoshi, 2016). This may be plausible  
66 since non-Gaussian data assimilation needs some information on higher-order moments of  
67 probability density functions (PDFs). As for the 4-dimensional variational method (4D-Var),  
68 Tsuyuki (2014) showed that the 4D-Var with a conventional cost function implicitly used a  
69 non-Gaussian prior PDF that evolved according to the Liouville equation (Ehrendorfer, 1994)  
70 if a certain condition was satisfied, and that the difficulty caused by multiple minima could  
71 be alleviated by combining with the EnKF. The iterative ensemble Kalman filter/smoothen  
72 (IEnKF/IEnKS) have been shown to be the missing link between the PF and the EnKF and

73 4D-Var, and can work very well with mild nonlinearity and generate a much better analysis  
74 than the above data assimilation methods (Sakov et al., 2012; Bocquet and Sakov, 2013,  
75 2014; Bocquet, 2016). However, the IEnKF/IEnKS need much larger computational cost due  
76 to the iterative application of the EnKF/EnKS and the use of a long assimilation window.

77       Recent developments in machine learning, in particular in deep learning (Le cum et al.,  
78 2015), have demonstrated impressive skills in various fields. Data-driven modeling,  
79 including data-driven parametrizations, based on machine learning has been extensively  
80 explored for improving simulations and predictions of nonlinear dynamical systems. Dueben  
81 and Bauer (2018) discussed the question of whether models that were based on deep  
82 learning and trained on atmospheric data could compete with weather and climate models  
83 that were based on physical principles. Reichstein et al. (2019) advocated a hybrid modeling  
84 approach in which physical process models were coupled with machine learning to further  
85 improve understanding and predictive ability in earth system science. Abarbanel et al. (2018)  
86 and Geer (2020) showed an equivalence in formulation between data assimilation and deep  
87 learning. Lists of literature of recent studies are available in Reichstein et al. (2019) and  
88 Chattopadhyay et al. (2020), for instance. Quite recently the combination of data assimilation  
89 and machine learning has been explored to address sparse and noisy observations in data-  
90 driven modeling (Brajard et al., 2020a; Bocquet et al., 2020; Tomizawa and Sawada, 2020;  
91 Gottwald and Reich, 2021; Wikner et al., 2021), data-driven parametrizations (Brajard et al.,  
92 2020b), and model error correction (Farchi et al., 2021).

93           Research on the application of deep learning to data assimilation itself has also started.  
94   Arcucci et al. (2021) proposed a method for integrating variational data assimilation with  
95   deep learning, in which a recurrent neural network is trained on the state of a dynamical  
96   model and the result of data assimilation. Silva et al. (2021) proposed the use of a generative  
97   adversarial network to make prediction and to assimilate observations by using a low-  
98   dimensional space for the spatial distribution of the simulated state. However, it is difficult to  
99   directly apply those methods to data assimilation in high-dimensional systems such as  
100   atmospheric models for numerical weather prediction.

101           In this study, we present an ensemble data assimilation method combining the EnKF  
102   and deep learning as an alternative to the PF for high-dimensional systems. The additional  
103   computational cost to assimilate observations is a very small fraction of that of the EnKF.  
104   Since a deep neural network (DNN) can learn a data assimilation method for a specific  
105   dynamical system and a specific observing system by training, we could expect this method  
106   to work with a relatively small ensemble size even in strongly nonlinear regimes. However,  
107   data assimilation in meteorology is generally a large-scale problem, and the background  
108   error covariance and the distribution of radar and satellite data change with the analysis time.  
109   The EnKF, as well as the PF and 4D-Var, can properly deal with this nonstationarity in data  
110   assimilation. On the other hand, deep learning is based on the minimization of the sum of  
111   errors over many samples. In addition, it would be difficult to provide sufficient information  
112   on the forecast error covariance to a DNN, because the feasible size of a DNN is limited,

113 where we define the size of a DNN as the total number of weights including bias parameters  
114 to be optimized by training. If the output of a DNN is not well optimized for each analysis  
115 time, the analysis accuracy may deteriorate in assimilation cycles. However, since the EnKF  
116 does not work very well in strongly nonlinear regimes, we could expect data assimilation by  
117 deep learning to outperform the EnKF in such regimes.

118 The purpose of this study is to propose a nonlinear data assimilation method based on  
119 deep learning that is locally embedded in an EnKF and to investigate its performance  
120 through data assimilation experiments in both perfect and imperfect model scenarios using  
121 toy models. By applying deep learning in combination with an EnKF, we can reduce the size  
122 of a DNN and address the nonstationarity in data assimilation. This method is named the  
123 deep learning-ensemble Kalman filter (DL-EnKF).

124 The remainder of this paper is organized as follows. Section 2 introduces the method of  
125 DL-EnKF. Section 3 describes the design of experiments in both perfect and imperfect model  
126 scenarios. Section 4 presents the results of these experiments. Summary and discussion  
127 are mentioned in Section 5.

128

## 129 **2. Method**

130 Since data assimilation is generally a large-scale problem, it is desirable to keep the  
131 size of a DNN as small as possible. For instance, the size of a feedforward neural network  
132 with  $m$  layers with  $n$  nodes per layer is about  $n^2(m-1)$  and a greater number of samples



133 would be required for training. If we directly apply deep learning to data assimilation, the  
134 number of input nodes is at least the sum of the number of observations and the degrees of  
135 freedom of a numerical model, and the number of output nodes is the degrees of freedom  
136 of the model, while the number of nodes of a hidden layer is usually required to be larger  
137 than the number of input or output nodes. For high-dimensional systems such as  
138 atmospheric models, the size of a DNN would become too large to be stored in the memory  
139 of a computer and to prepare sufficient training samples. To apply deep learning to data  
140 assimilation for atmospheric models, we need to introduce a localization procedure and to  
141 train a DNN to have some versatility so that it is applicable to each grid point in a certain  
142 range of geographical areas.

143 Figure 1a shows the workflow of the DL-EnKF, in which deep learning is locally  
144 embedded in an EnKF. The “EnKF” box in this figure represents the analysis step of the  
145 EnKF, and “Deep Learning” box consists of an ensemble of several DNNs. The inputs of  
146 DNNs to create the DL-EnKF analysis at a grid point are the EnKF analysis, forecast,  
147 observations, availability of observations in binary, and pseudo-observations that  
148 supplement missing observations. The EnKF analysis and forecast are the ensemble means  
149 of each ensemble. We do not explicitly use the information contained in the forecast  
150 ensemble other than the ensemble mean to reduce the size of DNNs. Since observational  
151 data for which the DNN has input nodes may be sometimes missing, it is necessary to  
152 provide the information on the availability of observations to DNNs. The pseudo-

Fig. 1

153 observations are prepared by using the EnKF analysis and the observation operators of the  
 154 missing observations. Those input data are extracted from a small domain centered at the  
 155 analysis grid point. The radius of this domain is hereafter referred to as the input radius, and  
 156 it is assumed that this value is smaller than the covariance localization radius of the EnKF.  
 157 According to Hsieh and Tang (1998), the DL-EnKF analysis is the average of outputs from  
 158 the ensemble of DNNs. The individual outputs from DNNs would be scattered in phase  
 159 space due to multiple minima of a loss function of deep learning, and we would likely obtain  
 160 a better DL-EnKF analysis by averaging those individual outputs.

161 The analysis ensemble  $\{\mathbf{x}_m^a\}_{m=1}^M$ , where  $M$  is the ensemble size, is created by  
 162 modifying the EnKF analysis ensemble  $\{\mathbf{x}_{\text{EnKF}, m}^a\}_{m=1}^M$  such that its ensemble mean is equal  
 163 to the DL-EnKF analysis  $\mathbf{x}_{\text{DL-EnKF}}^a$  as follows:

$$164 \quad \mathbf{x}_m^a = \mathbf{x}_{\text{DL-EnKF}}^a + \alpha(\mathbf{x}_{\text{EnKF}, m}^a - \mathbf{x}_{\text{EnKF}}^a), \quad (1)$$

165 where  $\mathbf{x}_{\text{EnKF}}^a$  is the EnKF analysis and  $\alpha$  is a parameter for adjusting the spread of the  
 166 analysis ensemble. If adaptive covariance inflation is used in the EnKF, we can set  $\alpha$  to 1  
 167 since the effect of tuning  $\alpha$  is almost canceled by this procedure. However, if we conduct  
 168 ensemble forecasts using the analysis ensemble, we may need to adjust the value of  $\alpha$ .  
 169 The members of the analysis ensemble thus generated are evolved by the time integration  
 170 of a numerical model to prepare the forecast ensemble for the next analysis time.

171 For the training of a DNN, we use the EnKF analysis and forecast provided by an EnKF  
 172 run as shown in Fig. 1b. The weights including bias parameters are optimized by reducing

173 a loss function that measures a difference between the output of the DNN and the target of  
174 training. We prepare several DNNs by randomly initializing the weights before the training.

175 One of the reasons for including the EnKF analysis in the inputs of DNNs is that this  
176 analysis at a grid point contains some information on the forecast, observations, and  
177 forecast error covariance in a domain within the covariance localization radius, so that we  
178 can reduce the input radius of DNNs and implicitly utilize some information of the forecast  
179 error covariance. In addition, even if DNNs cannot deal with some observational data  
180 because the input nodes for these observations are absent, they are assimilated by the  
181 EnKF part of DL-EnKF and their information is partly provided to the deep learning part  
182 through the EnKF analysis.

183 We can prepare pseudo-observations by other methods. However, it is easily shown by  
184 the sequential assimilation method (e.g., Houtekamer and Mitchell, 2001) that if observation  
185 errors are independent of each other the additional assimilation of pseudo-observations  
186 does not change the EnKF analysis. Therefore, it can be considered that the pseudo-  
187 observations thus created are assimilated in the EnKF part of DL-EnKF along with the real  
188 observations, and that the same observations including the pseudo-observations are  
189 provided to the deep learning part. In this sense, the method adopted in this study may be  
190 a natural choice, although it may not be optimal and spurious correlations and biases will be  
191 generated. We could expect that DNNs will learn to properly deal with this problem by  
192 training.

193        Lawson and Hansen (2004) showed that an analysis ensemble generated by a  
194        deterministic EnKF tends to retain multi-modality that may appear in a forecast ensemble,  
195        while this is not the case for a stochastic EnKF. Therefore, a stochastic EnKF is better than  
196        a deterministic EnKF for generating an analysis ensemble of the DL-EnKF. However, it is  
197        well known that if the ensemble size is relatively small, a stochastic EnKF is inferior in terms  
198        of the accuracy of the ensemble mean due to random perturbations that are added to  
199        observations (e.g., Sakov and Oke, 2008; Bowler et al., 2013), so that we adopt a  
200        deterministic EnKF for the EnKF part in the present paper.

201

## 202        **3. Design of experiments**

### 203        *3.1. Outline*

204        The performance of the DL-EnKF is investigated through both perfect and imperfect  
205        model experiments using three versions of the 40-variable Lorenz 96 models (Lorenz, 1996)  
206        and the serial ensemble square root filter (EnSRF; Whitaker and Hamill, 2002), which is one  
207        of the deterministic EnKFs. The ensemble size of the serial EnSRF is set to 10, because we  
208        are interested in the performance of the DL-EnKF with a relatively small ensemble. In this  
209        and the next sections, the EnKF means the serial EnSRF unless otherwise stated. The  
210        purpose of the perfect model experiments is to clarify the basic performance of the DL-EnKF,  
211        while that of the imperfect model experiments is to gain insight into the performance of the  
212        DL-EnKF when applied to data assimilation in the real atmosphere.

213 The experiments consist of two phases: the training phase of DNNs and the test phase  
214 using data assimilation experiments. In the training phase, we run the models and the EnKF  
215 to prepare training and validation datasets, which are used to train DNNs and to verify the  
216 accuracy of the output of DNNs, respectively. The length of period and time interval of these  
217 datasets are 1 000 and 1, respectively. This large time interval is taken to ensure that each  
218 data is almost independent of each other. In the test phase, we run the models to prepare a  
219 test dataset for the data assimilation experiments. The length of period of this dataset is also  
220 1 000. The accuracy of the DL-EnKF analysis is compared with the deep learning and EnKF  
221 analyses to evaluate the performance of the DL-EnKF. The workflow to create the deep  
222 learning analysis is the same as that of the DL-EnKF analysis except for the absence of  
223 feedback from the deep learning part to the EnKF part (Fig. 2). The analysis accuracy is  
224 evaluated by the RMSE that is the square root of the squared error averaged over the grid  
225 points and the period of the test dataset at a time interval of 1.

Fig. 2

226 In the perfect model experiments, we use the original 40-variable Lorenz 96 model and  
227 conduct two types of experiments, Exp-PA and Exp-PB, in which the targets of training are  
228 different. The target in Exp-PA is the true state generated by the model, while the target in  
229 Exp-PB is an analysis by the stochastic EnKF (Evensen, 1994; Burgers et al., 1998) with an  
230 ensemble size of 1 000. This analysis is hereafter referred to as the EnKF1000 analysis.  
231 Although this ensemble size may be unrealistic for the 40-variable model, the purpose of  
232 Exp-PB is to examine the performance of the DL-EnKF when an analysis with a high

233 accuracy is used as a target.

234 In the imperfect model experiments, the two-scale Lorenz 96 model with 40 large-scale  
235 variables and 400 small-scale variables is used as a substitute of the real atmosphere, while  
236 a parameterized Lorenz 96 model with a parameterization of large-scale forcing by small-  
237 scale variables is used as a substitute of a numerical model of the real atmosphere. We  
238 conduct two types of experiments, Exp-IA and Exp-IB. In Exp-IA, we train DNNs using the  
239 simulation data generated by the parameterized model, and conduct the data assimilation  
240 experiment using observations generated by the two-scale model. The idea behind Exp-IA  
241 is that if a dynamical system and a observing system that are used for the taining of a DNN  
242 resemble the real-world systems, we could expect that a data assimilation method the DNN  
243 has learned by training also works in the real-world applications. In Exp-IB, the two-scale  
244 Lorenz 96 model is used for the training and data assimilation experiments in a perfect model  
245 scenario for comparison to Exp-IA.

246 Table 1 summarizes the models used in the training and test phases of the experiments.

Table 1
---------

247 The following subsections describe further details of the experimental design.

248

### 249 3.2. Models

250 The governing equations of the Lorenz 96 model for the perfect model experiments are

$$251 \quad \frac{dX_k}{dt} = -X_{k-1}(X_{k-2} - X_{k+1}) - X_k + F, \quad (2)$$

252 for  $k = 1, \dots, K$ , satisfying periodic boundary conditions:  $X_{-1} = X_{K-1}$ ,  $X_0 = X_K$ , and  $X_1 =$

253  $X_{K+1}$ . The number of variables  $K$  and the forcing parameter  $F$  are set to 40 and 8,  
254 respectively. Note that since the number of positive Lyapunov exponents of the model is 13  
255 for those parameter values (Lorenz and Emanuel, 1998), the ensemble size of 10 is not very  
256 small. The leading Lyapunov exponent corresponds to a doubling time of 0.42 (Lorenz and  
257 Emanuel, 1998). When the nonlinearity in data assimilation is controlled by changing the  
258 time interval between observations as in the present study, this value can be used as a  
259 reference for estimating the degree of nonlinearity. The fourth-order Runge-Kutta scheme is  
260 adopted for the time integration of the model with a time step 0.01.

261 The governing equations of the two-scale Lorenz 96 model for the imperfect model  
262 experiments are

$$263 \quad \frac{dX_k}{dt} = -X_{k-1}(X_{k-2} - X_{k+1}) - X_k + F - \frac{hc}{b} \sum_{j=1}^J Y_{j,k}, \quad (3)$$

$$264 \quad \frac{dY_{j,k}}{dt} = -cbY_{j+1,k}(Y_{j+2,k} - Y_{j-1,k}) - cY_{j,k} + \frac{hc}{b} X_k, \quad (4)$$

265 for  $k = 1, \dots, K$  and  $j = 1, \dots, J$ , where  $\{X_k\}$  are large-scale variables and  $\{Y_{j,k}\}$  are small-  
266 scale variables, satisfying periodic boundary conditions:  $X_{-1} = X_{K-1}$ ,  $X_0 = X_K$ ,  $X_1 = X_{K+1}$ ,  
267  $Y_{0,1} = Y_{J,K}$ ,  $Y_{J+1,K} = Y_{1,1}$ , and  $Y_{J+2,K} = Y_{2,1}$ . To make Eq. (4) meaningful, we further define  
268  $Y_{0,k} = Y_{J,k-1}$ ,  $Y_{J+1,k} = Y_{1,k+1}$ , and  $Y_{J+2,k} = Y_{2,k+1}$ . Large- and small-scale variables interact  
269 with each other through the last terms on the right-hand side of Eqs. (3) and (4). We set the  
270 parameters as follows:  $K = 40$ ,  $J = 10$ ,  $F = 10$ ,  $h = 1$ ,  $c = 10$ , and  $b = 10$ . These values  
271 are the same as the ones used by Lorenz (1996) except for  $K$ . Note that the forcing  
272 parameter  $F$  is larger than in the perfect model experiments. The fourth-order Runge-Kutta

273 scheme is adopted for the time integration of the model with a time step 0.005.

274 We also need the parameterized Lorenz 96 model in the imperfect model experiments.  
275 Although advanced parametrization methods such as stochastic parametrization (e.g., Wilks,  
276 2005) and machine learning-based parametrization (e.g., Schneider et al., 2017) are  
277 available, a simple function fitting is adopted in the present study; the last term on the right-  
278 hand side of Eq. (3) is approximated by a linear function of  $X_k$ . The reason we adopt such  
279 a simple approach is that we intend to demonstrate that even an unsophisticated imperfect  
280 model works well for the training of a DNN. Then the governing equations of the  
281 parameterized Lorenz 96 model are

$$282 \quad \frac{dX_k}{dt} = -X_{k-1}(X_{k-2} - X_{k+1}) - X_k + F + (a_1 X_k + a_0), \quad (5)$$

283 for  $k = 1, \dots, K$ , satisfying periodic boundary conditions:  $X_{-1} = X_{K-1}$ ,  $X_0 = X_K$ , and  $X_1 =$   
284  $X_{K+1}$ . The number of variables  $K$  and the forcing parameter  $F$  were set to 40 and 10,  
285 respectively, to be consistent with the two-scale Lorenz 96 model. The constants  $a_1$  and  
286  $a_0$  are to be determined by the function fitting. The fourth-order Runge-Kutta scheme is  
287 adopted for the time integration of the model with a time step 0.01.

288 The three models are integrated from  $t = 0$  to  $t = 2\,050$  for preparing the training and  
289 validation datasets. The initial condition at each grid point is  $F$  plus an independent random  
290 number drawn from the normal distribution with the mean 0 and the variance 1, except that  
291 the small-scale variables of the two-scale Lorenz 96 model are set to 0 at the initial time.  
292 The data from  $t = 51$  to  $t = 1\,050$  are used for preparing the training dataset, and those



293 from  $t = 1\ 051$  to  $t = 2\ 050$  for the validation dataset. The other time integration of the  
294 models from  $t = 0$  to  $t = 1\ 050$  with initial conditions generated by using another random  
295 number sequence is conducted for preparing the test dataset, and the state variables from  
296  $t = 51$  to  $t = 1\ 050$  are used as the true state (target) for computing the analysis error.

297

### 298 **3.3. Observations**

299 Observations are generated by adding random errors to the results of the time  
300 integration of the models. The observation errors are independent random draws from the  
301 normal distribution with the mean 0 and the variance 1, so that the standard deviation of  
302 observation errors is 1. Observations used in the imperfect experiments are of large-scale  
303 variables of the two-scale Lorenz 96 model except for the training phase of Exp-IA, in which  
304 observations are prepared by the parameterized Lorenz 96 model.

305 Nonlinearity in data assimilation is controlled by changing the time interval between  
306 observations  $\Delta t$ . All experiments are performed for three values of  $\Delta t$ : 0.05, 0.20, and 0.50.  
307 The case of  $\Delta t = 0.05$  corresponds to a weakly nonlinear case, and that of  $\Delta t = 0.50$   
308 corresponds to a strongly nonlinear one. Note that the latter value is close to the doubling  
309 time 0.42 mentioned in Subsection 3.2, and that Penny and Miyoshi (2016) used  $\Delta t = 0.50$   
310 for their experiments of a local PF. All observations are prepared such that observations at  
311 the same analysis time are the same regardless of the time interval between observations.

312 For the spatial distribution of observations, two cases are examined. In one case,

313 observations are available at all grid points, and the number of observations is always 40.  
314 In other words, observations are available at each grid point with a probability of 1. In the  
315 other case, observations are available at each grid point with a probability of 1/2. It is  
316 assumed that events that an observation exists are independent of each other in space and  
317 time so that the spatial distribution of observations randomly changes at every observation  
318 time. The average number of observations is 20, and the standard deviation of the number  
319 of observations is  $\sqrt{40 \cdot (1/2)^2} \approx 3.16$ . Hence, the number of pseudo-observations used by  
320 deep learning is about the same as that of observations. The probability of observations is  
321 hereafter denoted by  $p$ .

322

### 323 *3.4. Data assimilation by EnKF*

324 Covariance localization and covariance inflation are needed to optimize the performance  
325 of the EnKF. The correlation function defined by Eq. (4.10) of Gaspari and Cohn (1999) is  
326 taken for covariance localization. The parameter  $c$  in this equation is regarded as the  
327 localization radius  $r_L$  (unit: grid intervals) in the present study, at which radius the  
328 correlation coefficient decreases to 5/24. An adaptive inflation method proposed by Li et al.  
329 (2009) is used for multiplicative covariance inflation. This method is based on the innovation  
330 statistics by Desroziers et al. (2005). Li et al. (2009) imposed lower and upper limits in the  
331 “observed” inflation factor  $\tilde{\Delta}^o$  before applying a smoothing procedure:  $0.9 \leq \tilde{\Delta}^o \leq 1.2$ . Since  
332 we conduct data assimilation over a much wider range of the time interval between

333 observations  $\Delta t$ , we optimize the upper limit of  $\tilde{\Delta}^o$  for each set of parameters  $(r_L, \Delta t, p)$   
334 leaving the lower limit at 0.9. The candidates of the upper limit are 1.2, 1.3, 1.4, 1.5, 2.0, 3.0,  
335 5.0, and no limit. In addition, although Li et al. (2009) set the error growth parameter  $\kappa$  to  
336 1.03, we adopt a larger value  $\kappa = 1.1$  because this value leads to a better analysis in the  
337 present study. A set of values of  $r_L$  and the upper limit of  $\tilde{\Delta}^o$  with the best analysis accuracy  
338 is hereafter referred to as the optimal parameters. We determine the optimal parameters for  
339 each pair of  $(\Delta t, p)$  in Exp-PA and Exp-IA by data assimilation experiments using the target  
340 and observations in each training dataset. The optimal parameters thus determined are also  
341 used in Exp-PB and Exp-IB, respectively, unless otherwise stated.

342 In Exp-PB, the target of training is the EnKF1000 analysis that is yielded by the  
343 stochastic EnKF with an ensemble size of 1 000, as mentioned in Subsection 3.1. The  
344 reason we adopt the stochastic EnKF is that when an ensemble size is very large the  
345 accuracy of the serial EnSRF tends to deteriorate and to become less accurate than the  
346 stochastic EnKF. We can avoid this problem with the serial EnSRF by applying the mean-  
347 preserving random rotation of an analysis ensemble (Sakov and Oke, 2008), but an  
348 additional computational cost is very large for the random rotation of a 1000-member  
349 ensemble. Figure 3 compares the analysis accuracy of the two EnKFs for ensemble sizes  
350 of 10 (in cold colors) and 1 000 (in warm colors), plotting the RMSEs averaged over the  
351 period from  $t = 51$  to  $t = 1050$ . This result is obtained by using the target and observations  
352 in the training datasets of Exp-PA. The localization radius and the upper limit of  $\tilde{\Delta}^o$  are

Fig. 3

353 optimized in the case of the ensemble size 10, while no covariance localization is applied  
354 and the upper limits of  $\tilde{\Delta}^o$  is set to 1.2 in the case of the ensemble size 1 000. It is found  
355 from Fig. 3 that in the latter case the RMSE of the stochastic EnKF is smaller than that of  
356 the serial EnSRF for all values of the time interval between observations  $\Delta t$  and probability  
357 of observations  $p$ . Note that the serial EnSRF with an ensemble size 10 outperforms the  
358 serial EnSRF with an ensemble size 1 000 in the three cases:  $(\Delta t, p) = (0.05, 1)$ ,  $(0.05, 1/2)$ ,  
359 and  $(0.20, 1)$ .

360 In Exp-IB, the two-scale Lorenz 96 model is used to assimilate observations of large-  
361 scale variables. As noted by Tsuyuki (2019), when the ensemble size is relatively small,  
362 forecast correlations between large- and small-scale variables are not reliable. Hence, these  
363 forecast correlations are neglected in the EnKF, and the analysis ensemble of small-scale  
364 variables is left unchanged from the forecast ensemble at each analysis time.

365

### 366 *3.5. Deep learning*

367 A simple feedforward neural network with the same number of nodes for all hidden layers  
368 is adopted for the deep learning part of the DL-EnKF. As we assume that the input radius of  
369 DNNs is relatively small, a convolutional neural network would not be needed. We could use  
370 a recurrent neural network instead of the feedforward neural network to utilize the  
371 information obtained by the previous processing in deep learning, but it is not adopted in the  
372 present study for simplicity.

373 The inputs of a DNN are the EnKF analysis using the optimal parameters, forecast,  
374 observations, availability of observations, and pseudo-observations in a small domain  
375 centered on an analysis grid point within the input radius  $r_l$  (unit: grid intervals). The  
376 availability of observation at a grid point is set to 1 if the observation is available and set to  
377 -1 if not available. The DNN assumes that observations are always available at all grid points,  
378 and we supplement missing observations with pseudo-observations. Since the availability  
379 of observations is not necessary in the case of  $p = 1$ , the input layer of the DNN has  $3(2r_l +$   
380  $1)$  nodes for  $p = 1$ , and  $4(2r_l + 1)$  nodes for  $p = 1/2$ . The number of hidden layers is set  
381 to 5 or 10 and the number of nodes per hidden layer is optimized as will be mentioned later.  
382 Since the balance of analysis (Kalnay, 2003) is not a serious issue in Lorenz 96 models, we  
383 let the output of the DNN be the analysis value at the analysis grid point only. For Exp-IB in  
384 which the two-scale Lorenz 96 model is used in the EnKF part of DL-EnKF, the input and  
385 output of the DNN are of large-scale variables only.

386 Table 2 summarizes the architecture and training of the DNN. All input and output data  
387 except for the availability of observations are normalized by using the mean and standard  
388 deviation of the target state in the training dataset of Exp-PA for the perfect model  
389 experiments and of Exp-IB for the imperfect model experiments. Since the statistical  
390 behavior of the models does not depend on the location of a grid point, the data at all grid  
391 points are used to prepare the training and validation datasets. Hence, the number of  
392 samples of each dataset is  $40 \times 1\,000 = 40\,000$ . The training dataset is split into small

Table 2
---------

393 batches called mini-batches that are used to compute the loss function and update the  
394 weights of the DNN. Learning rate decay is adopted in the training to avoid the situation in  
395 which the DNN converges towards minima in a noisy manner and ends up oscillating far  
396 away from actual minima. The number of epochs is the number of times each element in the  
397 training dataset is used by the DNN for optimizing the weights. For most of the cases,  
398 iterations almost converge within 10 epochs. We use PyTorch (Paszke et al., 2019) as the  
399 deep learning software.

400 To determine the optimal number of nodes per hidden layer, we train two DNNs with 5  
401 and 10 hidden layers by changing the number of nodes using the training and validation  
402 datasets of Exp-PA. Figure 4 plots the RMSEs of the two DNNs against the input radius  $r_I$   
403 for the case of  $\Delta t = 0.50$  and  $p = 1$ . The RMSEs are computed by using the validation  
404 datasets. For the DNN with 10 hidden layers and 5 nodes per hidden layer, the training fails  
405 for six values of the input radius, so that the RMSE of this case is not plotted in Fig. 4b.  
406 Since the RMSE is not very different between 5 and 10 hidden layers, we adopt the DNN  
407 with 5 hidden layers. It is found from Fig. 4 that when the input radius and number of nodes  
408 are large to some extent, the RMSE tends to increase due to the generalization error of deep  
409 learning. Since the DNN with 20 nodes per hidden layer has the smallest RMSE for most of  
410 the input radii, we set the optimal number of nodes to 20 for the case of  $\Delta t = 0.50$  and  $p =$   
411 1.

Fig. 4

412 The optimal numbers of nodes of the DNN with 5 hidden layers are summarized in Fig.

Fig. 5

413 5 by blue bars for  $p = 1$  and by cyan bars for  $p = 1/2$ . They tend to increase as the time  
414 interval between observations increases because the estimation of state variables becomes  
415 more difficult as nonlinearity increases. Although this result is obtained for Exp-PA, those  
416 number of nodes are used in all experiments. We also compute the optimal numbers of  
417 nodes for the case where the EnKF analysis is not included in the inputs of the DNN. The  
418 RMSEs for this case are plotted in Fig. 5 by red bars for  $p = 1$  and by orange bars for  $p =$   
419  $1/2$ . It is found that the inclusion of the EnKF analysis tends to reduce the optimal number  
420 of nodes. This is probably because the EnKF analysis plays the role of a first guess and  
421 makes it easier to estimate state variables.

422 The appendix discusses the impacts of the increase in the ensemble size of EnKF and  
423 the sample size for training on the accuracy of output of a DNN. The result of experiments  
424 shows that when the ensemble size of EnKF is increased to 40 the improvement by deep  
425 learning is considerably reduced, and that we need to increase the sample size much larger  
426 to obtain a larger improvement.

427

### 428 *3.6 Data assimilation by DL-EnKF*

429 In the data assimilation experiments with the DL-EnKF, the EnKF part of DL-EnKF  
430 adopts the optimal parameters, and the DL-EnKF analysis is the average of outputs of 5 or  
431 10 DNNs. Since the adaptive covariance inflation is used in the EnKF part, the parameter  
432  $\alpha$  in Eq. (1) is set to 1. In the test phase of Exp-IB, the deep learning part receives only

433 large-scale variables from the EnKF part and generates the DL-EnKF analysis of large-scale  
434 variables. The analysis ensemble of large-scale variables is modified by using this analysis,  
435 while the analysis ensemble of small-scale variables is left unchanged from the one  
436 generated by the EnKF part. The RMSEs of the DL-EnKF, deep learning, and EnKF analysis  
437 in Exp-IB are computed by using large-scale variables only.

438

## 439 **4. Results**

### 440 *4.1 perfect model experiments*

441 The first issue to be clarified is whether deep learning can outperform the EnKF in terms  
442 of analysis accuracy. The EnKF is close to optimal in weakly nonlinear regimes, and the  
443 deep learning part of DL-EnKF does not explicitly utilize the forecast error covariance. Figure Fig. 6  
444 6a compares the analysis accuracy between deep learning and the EnKF in Exp-PA for all  
445 values of  $\Delta t$  and  $p$ , in which the RMSEs are plotted against the RMSE of EnKF for the  
446 input radius of 2 grid intervals. The dots indicate the RMSEs of a single DNN, and the short  
447 horizontal bars indicate the RMSE of the average of outputs from 5 DNNs. It is found from  
448 this figure that all RMSEs of deep learning analysis are the same for each case and that  
449 deep learning outperforms the EnKF when  $\Delta t = 0.50$ . Note that since the EnKF analysis is  
450 included in the inputs of DNNs, the accuracy of deep learning analysis does not become  
451 worse than that of the EnKF analysis if sufficient training samples are available.

452 The second issue is whether the accuracy of the DL-EnKF analysis is better than that



453 of the EnKF and deep learning analyses. As mentioned in the introduction, the analysis by  
454 deep learning is based on the minimization of the sum of errors over many samples and not  
455 optimized for each analysis time. Hence, the analysis accuracy may deteriorate during  
456 assimilation cycles by the DL-EnKF. Figure 6b compares the analysis accuracy between the  
457 DL-EnKF and EnKF for Exp-PA. The dots indicate the RMSEs of DL-EnKF when the output  
458 of a single DNN is used as the DL-EnKF analysis, and the horizontal bars indicate the ones  
459 when the average of outputs from 5 DNNs is used as the DL-EnKF analysis. It is found that  
460 when  $\Delta t = 0.05$  the RMSEs of DL-EnKF based on a single DNN are scattered and larger  
461 than that of EnKF. In other words, the accuracy of the deep learning analysis shown in Fig.  
462 6a is not maintained during assimilation cycles in a weakly nonlinear case. Taking the  
463 average over 5 DNNs does not improve the accuracy very well. When  $\Delta t = 0.20$  and  $0.50$ ,  
464 on the other hand, the RMSEs of DL-EnKF based on a single DNN become almost the same  
465 for each case and taking the average over 5 DNNs slightly improves the accuracy in the  
466 case of  $\Delta t = 0.50$ . In addition, a comparison of Fig. 6a and 6b shows that the RMSE of DL-  
467 EnKF is smaller than that of deep learning when  $\Delta t = 0.50$  due to positive feedback in  
468 assimilation cycles.

469 We conduct additional experiments in which the ensemble size of DNNs is increased to  
470 10 in Exp-PA. The initial conditions of weights used for the training are different from the  
471 ones used in the case of 5 DNNs. Results are presented in Figs. 6c and 6d, and the former  
472 figure looks the same as Fig. 6a. The benefit of taking the average over 10 DNNs for the

473 DL-EnKF is evident when  $\Delta t = 0.05$ , although its analysis accuracy is still lower than that of  
474 EnKF. When  $\Delta t = 0.20$  and  $\Delta t = 0.50$ , the RMSEs of DL-EnKF are almost the same as in  
475 the case of 5 DNNs. This suggests that an ensemble size of 5 is sufficient except for a  
476 weakly nonlinear case. Then, all the results shown below are based on the average of  
477 outputs from 5 DNNs, because our interest is primarily in the performance of the DL-EnKF  
478 in strongly nonlinear regimes.

479 Figure 7 compares the time sequences of RMSEs of the EnKF (red line) and the DL-  
480 EnKF (green line) in the case of  $p = 1$ . When  $\Delta t = 0.05$  (Fig. 7a), the DL-EnKF is  
481 outperformed by the EnKF during the whole period. When  $\Delta t = 0.20$  (Fig. 7b), the analysis  
482 accuracy of the two methods is close; the correlation coefficient between the two RMSEs  
483 computed for the period from  $t = 51$  to  $t = 1\ 050$  is 0.761. When  $\Delta t = 0.50$  (Fig. 7c), the  
484 EnKF sometimes exhibits a significant deterioration of accuracy, but the DL-EnKF does not  
485 show such a tendency. This result demonstrates an excellent performance of the DL-EnKF  
486 in strongly nonlinear regimes.

Fig. 7

487 The third issue is whether the optimal input radius of deep learning is smaller than the  
488 optimal localization radius of the EnKF. Figure 8 plots the RMSEs of EnKF (orange lines),  
489 deep learning (green lines), and DL-EnKF (blue lines) analysis against the input radius for  
490 all cases of Exp-PA (solid lines) and Exp-PB (broken lines). The RMSE of EnKF in Exp-PB  
491 is the same as the one in Exp-PA. An orange broken line indicates the RMSE of the  
492 EnKF1000 analysis that is used for the training in Exp-PB. The optimal localization radius is

Fig. 8

493 plotted by a red arrow, except for the case of  $(\Delta t, p) = (0.05, 1/2)$  where the optimal  
494 localization radius is 11 grid intervals. The RMSEs of EnKF and deep learning overlap in  
495 Figs. 8a and 8b, and the RMSEs except for the EnKF1000 analysis almost overlap in Figs.  
496 8c and 8d.

497 When  $\Delta t = 0.05$  (Figs. 8a and 8b), the DL-EnKF is outperformed by the EnKF.  
498 Reflecting that an ensemble of 5 DNNs is not sufficient (see Fig. 5b), the graphs of the DL-  
499 EnKF are not smooth due to large sampling errors. When  $\Delta t = 0.20$  (Fig. 8c and 8d) the  
500 two data assimilation methods exhibit almost the same accuracy while when  $\Delta t = 0.50$   
501 (Figs. 8e and 8f) the DL-EnKF outperforms the EnKF irrespective of the input radius. The  
502 accuracy of the DL-EnKF analysis is higher than that of the deep learning analysis for the  
503 latter case due to positive feedback in assimilation cycles. We can conclude that the input  
504 radius of 2 grid intervals is sufficient to attain the best accuracy of the DL-EnKF analysis.  
505 This value is smaller than the optimal localization radii for both  $p = 1$  and  $p = 1/2$ . Even if  
506 the input radius is further increased, the accuracy of the DL-EnKF and deep learning  
507 analysis remains almost the same, although slight degradations are seen due to the  
508 generalization error of deep learning. This small sensitivity of RMSEs on the input radius  
509 indicates that the information at distant grid points contributes little to the DL-EnKF analysis  
510 even within the localization radius of the EnKF. The inclusion of the EnKF analysis in the  
511 inputs of DNNs also contributes to this insensitivity.

512 Another point to be noted in Figs. 8e and 8f is that even if DNNs are trained on the

513 EnKF1000 analysis, the accuracy of the DL-EnKF analysis is not very different from the one  
514 trained on the true state. Given the large errors of the EnKF1000 analysis shown in Figs. 8e  
515 and 8f, this result may look surprising. That is probably because this analysis well represents  
516 the basic dynamics of the Lorenz 96 model despite the large errors. If the difference in the  
517 ensemble size between the DL-EnKF and the target analysis is decreased, the accuracy of  
518 the DL-EnKF analysis in Exp-PB is more deteriorated. For instance, according to an  
519 additional experiment in which the ensemble size of the DL-EnKF is set to 40 for the case  
520 of  $\Delta t = 0.50$  and  $p = 1$  (see the appendix), the RMSEs of the EnKF analysis and DL-EnKF  
521 analyses in Exp-PA and Exp-PB are 0.682, 0.617, and 0.638, respectively, for the input  
522 radius of 2 grid intervals. The corresponding values for the ensemble size of 10 are 0.798,  
523 0.675, and 0.689 (see Fig. 8e), so that the deterioration of accuracy in Exp-PB is still not  
524 very large.

525 Finally, we examine the impact of including the EnKF analysis in the inputs of DNNs on  
526 the accuracy of the deep learning analysis using the test datasets of Exp-PA. Figure 9 plots  
527 the RMSE of deep learning in which the EnKF analysis is not included (cyan line) and the  
528 one in which the EnKF analysis is included (green line) against the input radius. The green  
529 lines are the same as in Fig. 8, and the two RMSEs overlap in Fig. 9f. For comparison, the  
530 RMSE of EnKF of which localization radius is not optimized is also plotted by an orange line  
531 against the localization radius with the upper limit of  $\tilde{\Delta}^o$  optimized for each localization  
532 radius. Note that the RMSE of EnKF does not always attain the minimum at the optimal

Fig. 9

533 localization radius indicated by a red arrow, because the values of the optimal parameters  
534 are determined by using the training datasets.

535 We can see from Fig. 9 that the accuracy of the deep learning analysis is improved by  
536 including the EnKF analysis in the inputs of DNNs except for Fig. 9f, in which the EnKF  
537 analysis is too inaccurate to be useful. It is also found in Figs. 9c and 9e for  $p = 1$  that this  
538 procedure reduces the dependence of the analysis accuracy on the input radius. This is  
539 because the EnKF analysis contains some information on the forecast ensemble and  
540 observations in a domain within the localization radius. Such a reduction in the dependence  
541 brought about by including the EnKF analysis is not clearly seen in Figs. 9d and 9f for  $p =$   
542  $1/2$ , since deep learning partly utilizes the EnKF analysis through pseudo-observations.

543

#### 544 *4.2 Imperfect model experiments*

545 The parametrization procedure for the parameterized Lorenz 96 model is described in  
546 Subsection 3.2. Figure 10 is the scatter plot between the large-scale variables and the Fig. 10  
547 forcing. The initial condition is the same as that used for preparing the training dataset of  
548 Exp-IB. The number of samples is 40 000 and the result of linear function fitting is plotted by  
549 a straight line. The values of constants in Eq. (5) are  $a_1 = -0.320$  and  $a_0 = -0.165$ . Since  
550 the slope of this line is negative, the forcing acts on large-scale variables as negative  
551 feedback. Figure 11 compares the Hovmöller diagrams of the Lorenz 96 model, Fig. 11  
552 parameterized Lorenz 96 model, and large-scale variables of the two-scale Lorenz 96 model.

553 The initial condition is the same as that used in Fig. 10. Note that the forcing parameter  $F$   
554 is larger than in the perfect model experiments. A comparison of the three panels in Fig. 11  
555 shows that the parameterization works well, although the parametrized Lorenz 96 model  
556 evolves a little more regularly than the two-scale Lorenz 96 model. Stochastic  
557 parameterizations could remedy this defect (Wilks, 2005).

558 Figure 12 plots the RMSEs of EnKF (orange lines), deep learning (green lines), and DL-  
559 EnKF (blue lines) analysis against the input radius for all cases of Exp-IA (solid lines) and  
560 Exp-IB (broken lines). Note that Exp-IA is conducted in an imperfect model scenario, while  
561 Exp-IB is conducted in a perfect model scenario for comparison. Unlike Fig. 8, an orange  
562 broken line indicates the RMSE of EnKF using the two-scale Lorenz 96 model. The optimal  
563 localization radius of the EnKF is indicated by a red arrow. The RMSEs of EnKF and deep  
564 learning for each case overlap in Figs. 12a-12d. It is found that the RMSE of EnKF using  
565 the two-scale Lorenz 96 model is smaller than the one using the parameterized Lorenz 96  
566 model. We can confirm that the basic performance of the DL-EnKF is the same as in the  
567 perfect model experiments; the DL-EnKF is inferior to the EnKF in a weakly nonlinear case  
568 (Figs. 12a and 12b), while the opposite is true in a strongly nonlinear case (Figs. 12e and  
569 12f), in which the optimum input radius is smaller than the optimum localization length. A  
570 difference from the perfect model experiments is that when  $\Delta t = 0.20$  (Figs. 12c and 12d),  
571 the accuracy of the DL-EnKF analysis is a little worse than that of the EnKF analysis for  $p =$   
572 1 and a little better for small values of the input radius for  $p = 1/2$ .

Fig. 12

573 An important point to be noted in Figs. 12e and 12f is that even if DNNs are trained on  
574 the training dataset prepared by the parameterized Lorenz 96 model, the improvement in  
575 analysis accuracy introduced by the DL-EnKF is not very different from the case where the  
576 training dataset is prepared by the two-scale Lorenz 96 model. The former model is run in  
577 the data assimilation experiments in the test phase of Exp-IA without any trouble, implying  
578 that this model well represents the basic dynamics of large-scale variables of the two-scale  
579 Lorenz 96 model. When we use the Lorenz 96 model with  $F = 10$ , of which evolution is  
580 shown in Fig.11a, in the above data assimilation experiments, we often experience failures.

581

## 582 **5. Summary and discussion**

583 An ensemble data assimilation method based on deep learning was presented, in which  
584 an ensemble of DNNs was locally embedded in an EnKF. This method was named the DL-  
585 EnKF. The inputs of a DNN were the EnKF analysis, forecast, observations, availability of  
586 observations, and pseudo-observations in a small domain centered on an analysis grid point.  
587 Missing observations were supplemented with the pseudo-observations created from the  
588 EnKF analysis. The DL-EnKF analysis was the average of outputs from an ensemble of  
589 DNNs. The DL-EnKF analysis ensemble was generated from the DL-EnKF analysis and the  
590 EnKF analysis deviation ensemble. The members of the DL-EnKF analysis ensemble thus  
591 generated were evolved by the time integration of a numerical model to prepare the forecast  
592 ensemble for the next analysis time.

593 The performance of the DL-EnKF was investigated through data assimilation  
594 experiments in both perfect and imperfect model scenarios using three versions of the  
595 Lorenz 96 model and the serial EnSRF with an ensemble size of 10. The target of training  
596 in the perfect model experiments was the true state generated by the Lorenz 96 model or  
597 the EnKF1000 analysis generated by the stochastic EnKF with an ensemble size of 1000.  
598 In the imperfect model experiments, the true state and observations were provided by the  
599 two-scale Lorenz 96 model, while the training dataset was prepared by using the  
600 parameterized Lorenz 96 model. Nonlinearity in data assimilation was controlled by  
601 changing the time interval between observations.

602 The DL-EnKF was outperformed by the serial EnSRF in a weakly nonlinear case, but it  
603 was superior to the serial EnSRF in terms of analysis accuracy in a strongly nonlinear case  
604 despite such a small ensemble size. The DL-EnKF analysis was more accurate than the  
605 output of deep learning due to positive feedback in assimilation cycles in the latter case.  
606 Even if the target of training was the EnKF1000 analysis or the simulation by the  
607 parametrized Lorenz 96 model, the improvement introduced by the DL-EnKF was not very  
608 different from the case where the target of training was the true state. The inclusion of EnKF  
609 analysis in the inputs of DNNs not only improved the accuracy of the deep learning analysis  
610 but also reduced the optimal number of nodes per hidden layer and the dependence of the  
611 accuracy on the input radius.

612 Although the above results were obtained from experiments using toy models, they



613 suggest that the DL-EnKF may be a promising methods for data assimilation in strongly  
614 nonlinear regimes. The DL-EnKF works with a relatively small ensemble size compared to  
615 the PF, and we can prepare a training dataset for deep learning from simulation data by a  
616 numerical model used in data assimilation. Observational data and EnKF analysis data  
617 generated with a large ensemble could be used for this purpose, but a huge computational  
618 cost may be needed to obtain sufficient samples and a period when observations are  
619 available is limited.

620 The DL-EnKF may be suitable for data assimilation in cloudy or convective regions in  
621 the atmosphere to assimilate radar and satellite observations. We need to extend the inputs  
622 and output of DNNs in the vertical to assimilate satellite radiance data, since they are  
623 nonlocal observations. As for radial wind data by a Doppler radar, the direction and distance  
624 of a radar site differ depending on a grid point. However, if the radar site position relative to  
625 the grid point is included in the inputs of a DNN, we can train the DNN collectively regardless  
626 of the grid point as in the present study.

627 There are a couple of issues to be addressed before applying the DL-EnKF to data  
628 assimilation in the atmosphere. In the data assimilation experiments using Lorenz 96 models,  
629 the analysis value at a single grid point is sufficient for the output of a DNN, but we need to  
630 take the balance of analysis into account for atmospheric data assimilation. One of the  
631 methods for ensuring the balance is to extend the output of a DNN to include analysis values  
632 at surrounding grid points. Then the target of training consists of the target state in a small

633 domain centered on an analysis grid point. Since the target state is usually well balanced,  
634 the DNN could learn the balance. Adding a penalty term for suppressing imbalance to a loss  
635 function of the DNN may help enhance the balance. In addition, taking a weighting average  
636 of the outputs in adjacent domains may be effective in improving analysis accuracy.

637 This study demonstrates that the DL-EnKF is inferior to the EnKF in a weakly nonlinear  
638 case. It is found that an increase in the ensemble size of DNNs can mitigate this problem,  
639 but it would be difficult to increase the ensemble size sufficiently, given the computational  
640 cost needed for the training of DNNs. We may need a criterion for replacing the EnKF  
641 analysis with the corresponding DL-EnKF analysis in DL-EnKF assimilation cycles. An  
642 advanced DNN such as a recurrent neural network would be useful for improving the  
643 performance of DL-EnKF in weakly nonlinear regimes as well as in strongly nonlinear ones.

644

#### 645 **Data Availability Statement**

646 The Python programs used for Exp-PA in this study are available on J-STAGE Data.

647

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657

658

## Appendix

659 In this study, we perform the experiments using the serial EnSRF with 10 members and  
660 the 40 000 training samples. It may be of interest to examine how the accuracy of output of  
661 a DNN changes when the ensemble size and the sample size are increased. This appendix  
662 presents some results of additional experiments in which the ensemble size of EnKF is set  
663 to 10 and 40 and the sample size is set to 40 000, 160 000 and 640 000. These experiments  
664 correspond to Exp-PA for  $\Delta t = 0.50$ . The ensemble size of 40 is the same as the degrees  
665 of freedom of the Lorenz 96 model and, according to Fig. 5 of Penny and Miyoshi (2016),  
666 an EnKF still outperforms a local PF with this ensemble size. The periods of time integration  
667 of the model for preparing the training and validation datasets are 2 050, 8 050, and 32 050  
668 with the first periods of 50 in length are discarded.

669 Figure A1 shows the optimum numbers of nodes per hidden layer of a DNN, obtained  
670 by using the validation datasets. The maximum number of nodes is limited to 100. Although  
671 we choose the number of nodes that performs the best for the various input radius, the  
672 determination of the optimal number becomes difficult with the increase of the sample size.

Fig. A1

673 When the number of training samples is increased, the generalization error of deep learning  
674 tends to reduce and, therefore, the optimal number of nodes per hidden layer tends to  
675 increase.

676 Figure A2 compares RMSE between the serial EnSRF and the output of a DNN obtained  
677 by using the test datasets of Exp-PA. Note that they are not the average over 5 DNNs, so  
678 that the RMSEs shown by green lines in Figs. A2a and A2b are different from those in Figs.  
679 8e and 8f, respectively. The optimal localization radius of the serial EnSRF in Fig. A2c is 12  
680 grid intervals. It is found from this figure that the improvement by deep learning is  
681 considerably reduced for the ensemble size of 40. When the sample size is increased, the  
682 RMSE of the output of a DNN is reduced, but we need much more training samples to obtain  
683 a large improvement.

Fig. A2

684

685

## References

686

687 Abarbanel, H. D., P. J. Rozdeba, and S. Shirman, 2018: Machine learning: deepest learning  
688 as statistical data assimilation problems. *Neural Computation*, **30**, 2025–2055.

689 Arcucci, R., J. Zhu, S. Hu and Y.-K. Guo, 2021: Deep data assimilation: Integrating deep  
690 learning with data assimilation. *Appl. Sci.*, **11**, 1114-1134. doi:10.3390/app11031114.

691 Bocquet, M., 2016: Localization and the iterative ensemble Kalman smoother. *Quart. J. Roy.*  
692 *Meteor. Soc.*, **142**, 1075–1089. doi:10.1002/qj.2711.

- 693 Bocquet, M., J. Brajard, A. Carrassi and L. Bertino, 2020: Bayesian inference of chaotic  
694 dynamics by merging data assimilation, machine learning and expectation-maximization.  
695 *Foundations of Data Science*, **2**, 55–80.
- 696 Bocquet, M., C. A. Pires and L. Wu, 2010: Beyond Gaussian statistical modeling in  
697 geophysical data assimilation. *Mon. Wea. Rev.*, **138**, 1997-3023.  
698 doi:10.1175/2010MWR3164.1.
- 699 Bocquet, M. and P. Sakov, 2013: Joint state and parameter estimation with an iterative  
700 ensemble Kalman smoother. *Nonlin. Processes Geophys.*, **20**, 803–818.  
701 doi:10.5194/npg-20-803-2013.
- 702 Bocquet, M. and P. Sakov, 2014: An iterative ensemble Kalman smoother. *Quart. J. Roy.*  
703 *Meteor. Soc.*, **140**, 1521–1535. doi:10.1002/qj.2236.
- 704 Bowler, N. E., J. Flowerdew, and S. R. Pring, 2013: Tests of different flavours of EnKF on a  
705 simple model. *Quart. J. Roy. Meteor. Soc.*, **139**, 1505-1519. doi:10.1002/qj.2055.
- 706 Brajard, J., A. Carrassi, M. Bocquet and L. Bertino, 2020a: Combining data assimilation and  
707 machine learning to emulate a dynamical model from sparse and noisy observations: a  
708 case study with the Lorenz 96 model. *J. Comp. Sci.*, **44**, 101171.
- 709 Brajard, J., A. Carrassi, M. Bocquet and L. Bertino, 2020b: Combining data assimilation and  
710 machine learning to infer unresolved scale parametrisation. *Philos. Trans. Roy. Soc.*  
711 *London, Ser. A*, **379**, 2020.0086. doi:10.1098/rsta.2020.0086.
- 712 Burgers, G., P. J. van Leeuwen, and G. Evensen, 1998: Analysis schemes in the ensemble

713 Kalman filter. *Mon. Wea. Rev.*, **126**, 1719-1724.

714 Chattopadhyay, A., P. Hassanzadeh and D. Subramanian, 2020: Data-driven prediction of a  
715 multi-scale Lorenz 96 chaotic system using deep learning methods: Reservoir  
716 computing, artificial neural network, and long short-term memory network. *Nonlinear  
717 Processes Geophys.*, **27**, 373–389. doi:10.5194/npg-27-373-2020.

718 Desroziers, G., L. Berre, B. Chapnik, and P. Poli, 2005: Diagnosis of observation,  
719 background and analysis-error statistics in observation space. *Quart. J. Roy. Meteor.  
720 Soc.*, **131**, 3385–3396.

721 Dueben, P. D. and P. Bauer, 2018: Challenges and design choices for global weather and  
722 climate models based on machine learning, *Geosci. Model Dev.*, **11**, 3999–4009.  
723 doi:10.5194/gmd-11-3999-2018.

724 Ehrendorfer, M., 1994: The Liouville equation and its potential usefulness for the prediction  
725 of forecast skill. Part I: Theory. *Mon. Wea. Rev.*, **122**, 703-713.

726 Evensen, G., 1994: Sequential data assimilation with a nonlinear quasigeostrophic model  
727 using Monte Carlo methods to forecast error statistics. *J. Geophys. Res.*, **99**, 10143-  
728 10162.

729 Farchi, A. and M. Bocquet, 2018: Review article: Comparison of local particle filters and new  
730 implementations, *Nonlinear Processes Geophys.*, **25**, 765-807. doi:10.5194/npg-25-  
731 765-2018.

732 Farchi, A. M. Laloyaux, M. Bonavita and M. Bocquet, 2021: Using machine learning to correct

733 model error in data assimilation and forecast applications. *Quart. J. Roy. Meteor. Soc.*,  
734 **147**, 3067-3084. doi:10.1002/qj.4116.

735 Gaspari, G. and S. E. Cohn, 1999: Construction of correlation functions in two and three  
736 dimensions. *Quart. J. Roy. Meteor. Soc.*, **125**, 723-757.

737 Geer, A. J., 2020: Learning earth system models from observations: machine learning or  
738 data assimilation? *Technical Memorandum 863*, ECMWF, 23pp.

739 Gottwald, G. A. and Reich, S. (2021). Supervised learning from noisy observations:  
740 Combining machine-learning techniques with data assimilation. *Physica D: Nonlinear  
741 Phenomena*, **423**, 132911. doi:10.1016/j.physd.2021.132911.

742 Houtekamer, P. L. and H. L. Mitchell, 2001: A sequential ensemble Kalman filter for  
743 atmospheric data assimilation. *Mon. Wea. Rev.*, **129**, 123-137.

744 Hsieh, W. W. and B. Tang, 1998: Applying neural network models to prediction and data  
745 analysis in meteorology and oceanography. *Bull. Amer. Meteor. Soc.*, **79**, 1855-1870.

746 Kalnay, E., 2003: *Atmospheric Modeling, Data Assimilation and Predictability*. Cambridge  
747 University Press, Cambridge, 364 pp.

748 Kawabata, T. and G. Ueno, 2020: Non-Gaussian probability densities of convection initiation  
749 and development investigated using a particle filter with a storm-scale numerical  
750 weather prediction model. *Mon. Wea. Rev.*, **148**, 3–20. doi:10.1175/MWR-D-18-0367.1.

751 Kingma, D. and J. Ba, 2015: Adam: A method for stochastic optimization. arXiv:1412.6980.

752 Kondo, K. and T. Miyoshi, 2019: Non-Gaussian statistics in global atmospheric dynamics: a

753 study with a 10 240-member ensemble Kalman filter using an intermediate atmospheric  
754 general circulation model. *Nonlinear Processes Geophys.*, **26**, 211–225.  
755 doi:10.5194/npg-26-211-2019.

756 Lawson, G. W. and J. A. Hansen, 2004: Implications of stochastic and deterministic filters  
757 as ensemble-based data assimilation methods in varying regimes of error growth. *Mon.*  
758 *Wea. Rev.*, **132**, 1966-1981.

759 Le Cum, Y., Y. Bengio and G. Hinton, 2016: Deep learning. *Nature*, **521**, 436-444.  
760 doi:10.1038/nature14539.

761 Li, H., E. Kalnay and T. Miyoshi, 2009: Simultaneous estimation of covariance inflation and  
762 observation errors within an ensemble Kalman filter. *Quart. J. Roy. Meteor. Soc.*, **135**,  
763 523-533. doi:10.1002/qj.371.

764 Lorenz, E. D., 1996: Predictability – A problem partly solved, *Proceedings of the ECMWF*  
765 *Seminar on Predictability (4-9 September 1995, Reading, UK)*, ECMWF, 1-18.

766 Lorenz, E. D. and K. A. Emanuel, 1998: Optimal sites for supplementary weather  
767 observations: Simulation with a small model. *J. Atmos. Sci.*, **55**, 399-414.

768 Paszke, A., S. Gross, F. Massa, A. Lerer, J. Bradbury, G. Chanan, T. Killeen, Z. Lin, N.  
769 Gimelshein, L. Antiga, A. Desmaison, A. Köpf, E. Yang, Z. DeVito, M. Raison, A.  
770 Tejani, S. Chilamkurthy, B. Steiner, L. Fang, J. Bai and S. Chintala, 2019: PyTorch: An  
771 imperative style, high-performance deep learning library. In H. Wallach et al.,  
772 eds. *Advances in Neural Information Processing Systems 32*. Curran Associates, Inc.,



773 8024–8035. Available at: <http://papers.neurips.cc/paper/9015-pytorch-an-imperative->  
774 [style-high-performance-deep-learning-library.pdf](http://papers.neurips.cc/paper/9015-pytorch-an-imperative-style-high-performance-deep-learning-library.pdf).

775 Penny, S. G. and T. Miyoshi, 2016: A local particle filter for high-dimensional geophysical  
776 systems, *Nonlinear Processes Geophys.*, **23**, 391-405. doi:10.5194/npg-23-391-2016.

777 Reichstein, M., G. Camps-Vallis, B. Stevens, M. Jung, J. Denzler and N. Carvalhais and  
778 Prabhat, 2019: Deep learning and process understanding for data-driven Earth system  
779 science, *Nature*, **566**, 195–204. doi:10.1038/s41586-019-0912-1.

780 Sakov, P. and P. R. Oke, 2008: Implications of the form of the ensemble transformation in  
781 the ensemble square root filters. *Mon. Wea. Rev.*, **136**, 1042–1053.  
782 doi:10.1175/2007MWR2021.1.

783 Sakov, P., D. S. Oliver and L. Bertino, 2012: An iterative EnKF for strongly nonlinear systems.  
784 *Mon. Wea. Rev.*, **140**, 1988–2004. doi:10.1175/MWR-D-11-00176.1.

785 Schneider, T., S. Lan, A. Stuart and J. Teixeira, 2017: Earth system modeling 2.0: A blueprint  
786 for models that learn from observations and targeted high-resolution simulations.  
787 *Geophys. Res. Lett.*, **44**, 12396–12417. doi:10.1002/2017GL076101.

788 Silva, V. L., C. E. Heaney, Y. Li and C. C. Pain, 2021: Data Assimilation Predictive GAN  
789 (DAPredGAN): Applied to determine the spread of COVID-19. arXiv:2105.07729.

790 Snyder, C, T. Bengtsson, P. Bickel and J. Anderson, 2008: Obstacles to high dimensional  
791 particle filtering. *Mon. Wea. Rev.* **136**, 4629–4640. doi:10.1175/2008MWR2529.1

792 Tomizawa, F. and Y. Sawada, 2020: Combining ensemble Kalman filter and reservoir

793 computing to predict spatio-temporal chaotic systems from imperfect observations and  
794 models. *Geosci. Model Dev. Discuss.* doi:10.5194/gmd-2020-211.

795 Tsuyuki, T., 2014: Deterministic predictability of the most probable state and reformulation  
796 of variational data assimilation. *J. Meteor. Soc. Japan*, **92**, 599-622.  
797 doi:10.2151/jmsj.2014-606.

798 Tsuyuki, T., 2019: Ensemble Kalman filtering based on potential vorticity for atmospheric  
799 multi-scale data assimilation. *J. Meteor. Soc. Japan*, **97**, 1191-1210.  
800 doi:10.2151/jmsj.2019-067.

801 Whitaker, J. S. and T. M. Hamill, 2002: Ensemble data assimilation without perturbed  
802 observations. *Mon. Wea. Rev.*, **130**, 1913-1924.

803 Wikner, A., J. Pathak, B. R. Hunt, I. Szunyogh, M. Girvan and E. Ott, 2021: Using data  
804 assimilation to train a hybrid forecast system that combines machine-learning and  
805 knowledge-based components. *Chaos: An Interdisciplinary Journal of Nonlinear  
806 Science*, **31**, 053114.

807 Wilks, D. J., 2005: Effects of stochastic parametrizations in the Lorenz '96 system. *Quart. J.  
808 Roy. Meteor. Soc.*, **131**, 389-407. doi:10.1256/qj.04.03.

809 van Leeuwen, P. J., 2009: Particle filtering in geophysical systems. *Mon. Wea. Rev.*, **137**,  
810 4089-4114. doi:10.1175/2009MWR2835.1.

811 van Leeuwen, P. J., H. R. Kunsch, L. Nerger, R. Potthast and S. Reich, 2019: Particle filters  
812 for high-dimensional geoscience applications: A review. *Quart. J. Roy. Meteor. Soc.*, **145**,

813 2335-2365. doi:10.1002/qj.3551.

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819

820 Fig. 2 Workflow to generate deep learning analysis.

821

822 Fig. 3 Comparison of RMSEs between the serial EnSRF (abscissa) and stochastic EnKF  
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829 Fig. 4 Comparison of RMSEs of a single DNN with 5 (light green), 10 (orange), 20 (red),  
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833 is 1.

834

835 Fig. 5 The optimal number of nodes per hidden layer of a DNN with 5 hidden layers. The  
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845 indicate RMSEs based on a single DNN, and short horizontal bars indicate RMSEs based  
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848

849 Fig. 7 Time sequences of RMSEs of EnKF (red lines) and DL-EnKF (blue lines) for  
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860

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864 abscissa is the input radius for deep learning and the localization radius for the EnKF. A  
865 red arrow indicates the optimal localization radius. The observation time interval and the  
866 probability of observations are (a) 0.05 and 1, (b) 0.05 and 1/2, (c) 0.20 and 1, (d) 0.20  
867 and 1/2, (e) 0.50 and 1, (f) 0.50 and 1/2, respectively.

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869 Fig. 10 Scatter plot between large-scale variables (abscissa) and large-scale forcing by  
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872

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875

876 Fig. 12 Same as Fig. 8 except for Exp-IA (solid lines) and Exp-IB (broken lines) and that  
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878

879 Fig. A1 The optimal number of nodes per hidden layer of a DNN with 5 hidden layers for  
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882 observations 1 and 1/2, respectively. Red and orange bars are for the EnKF ensemble  
883 size of 40 for the probability of observations 1 and 1/2, respectively.

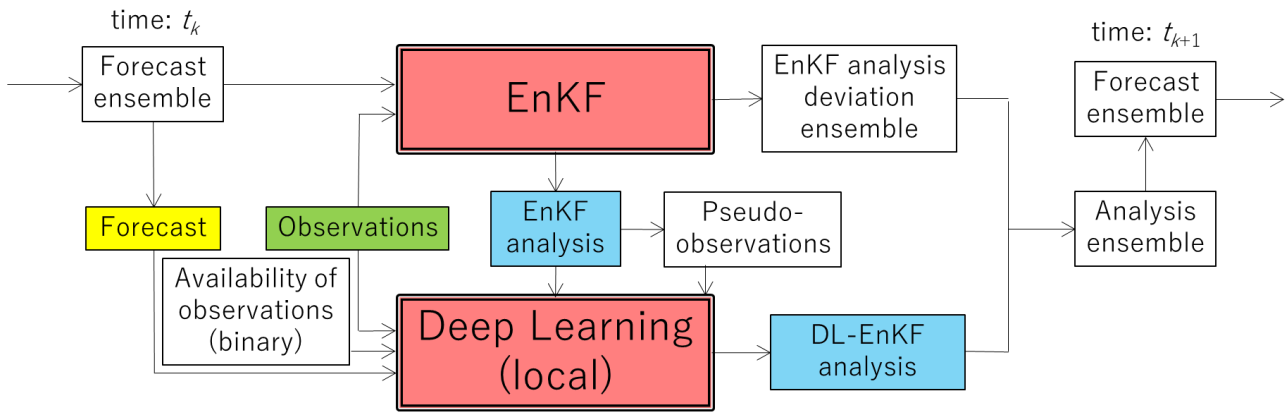
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885 Fig. A2 Comparison of RMSE between EnKF (orange lines) and the output of a DNN with  
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888 observations are (a) 10 and 1, (b) 10 and 1/2, (c) 40 and 1, (d) 40 and 1/2, respectively.  
889 The RMSEs are plotted against the input radius, and a red arrow indicates the optimal  
890 localization radius of EnKF.

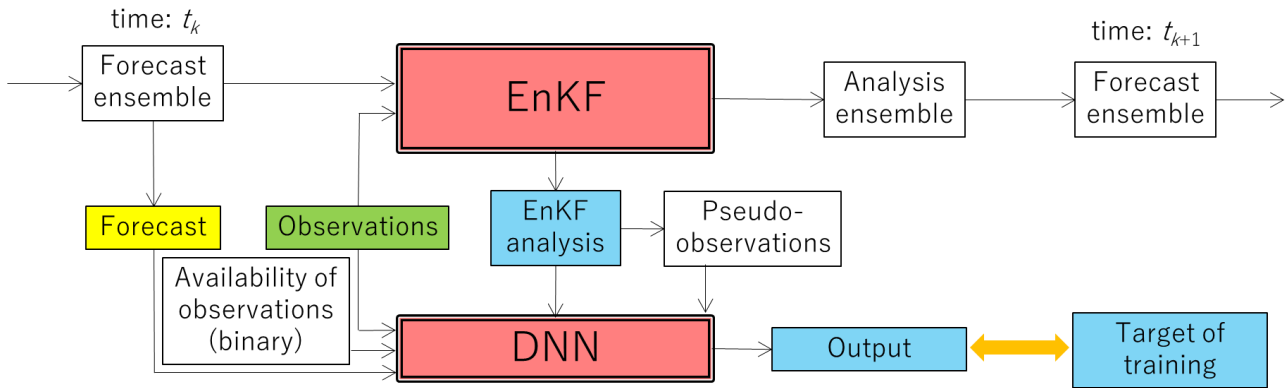
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893 (a)



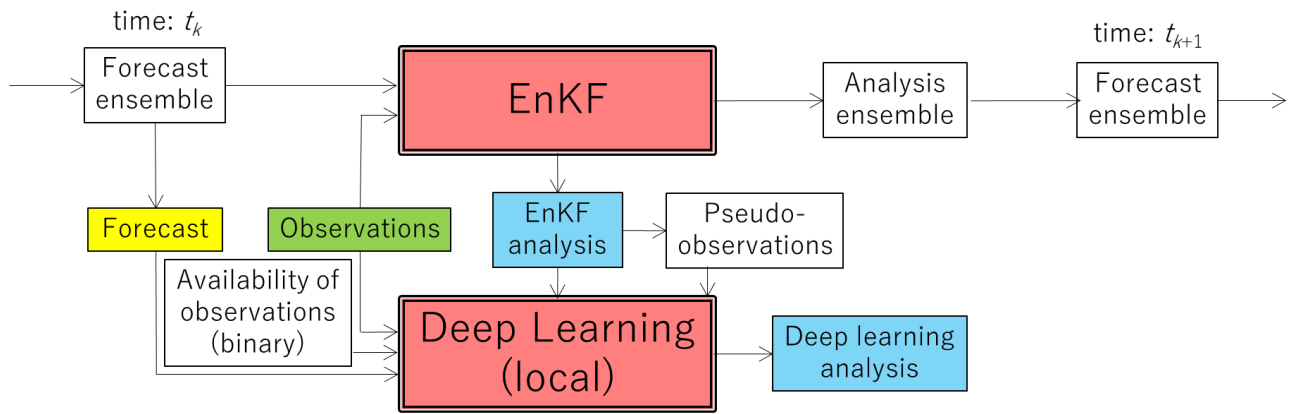
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895 (b)



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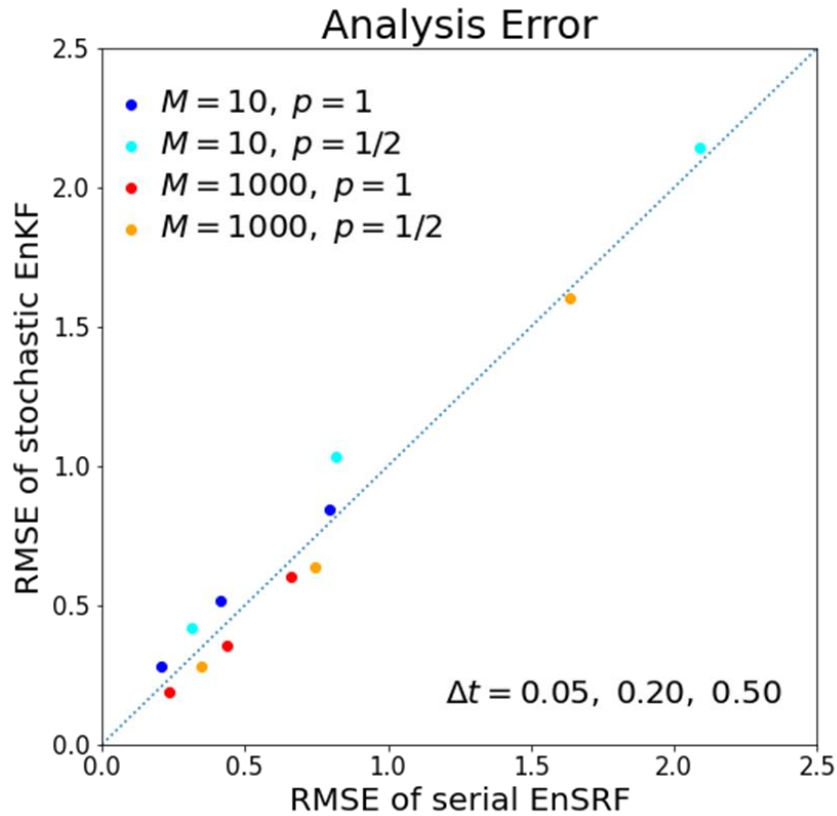




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901 Fig. 2 Workflow to generate deep learning analysis.

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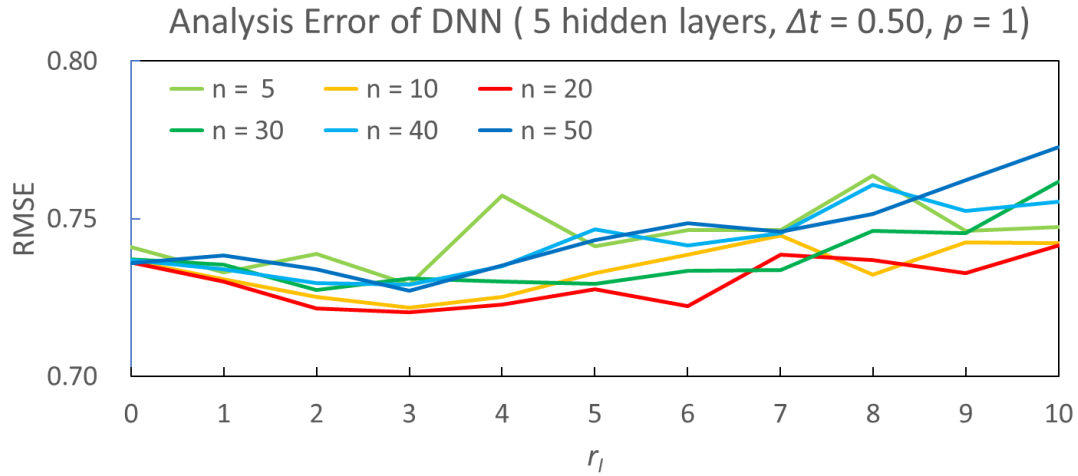


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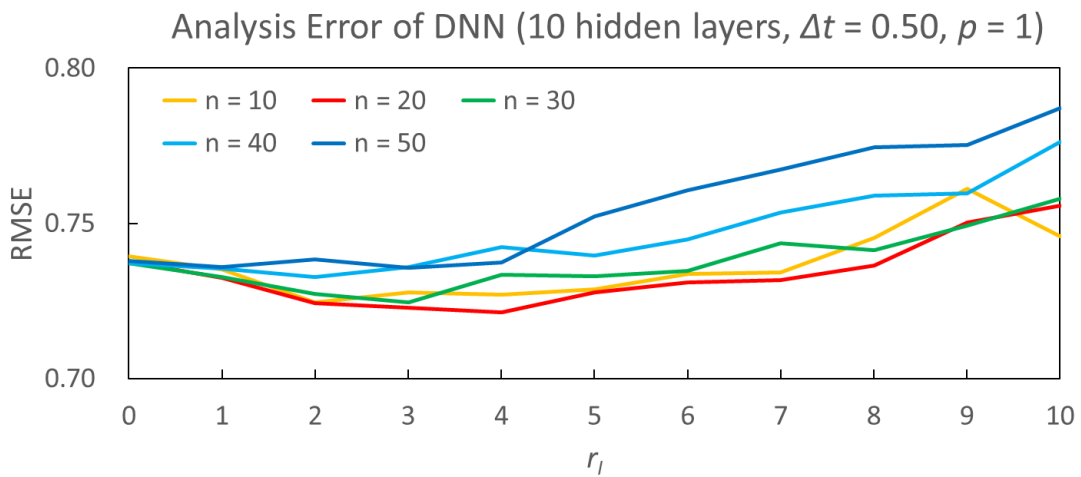
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912 (a)



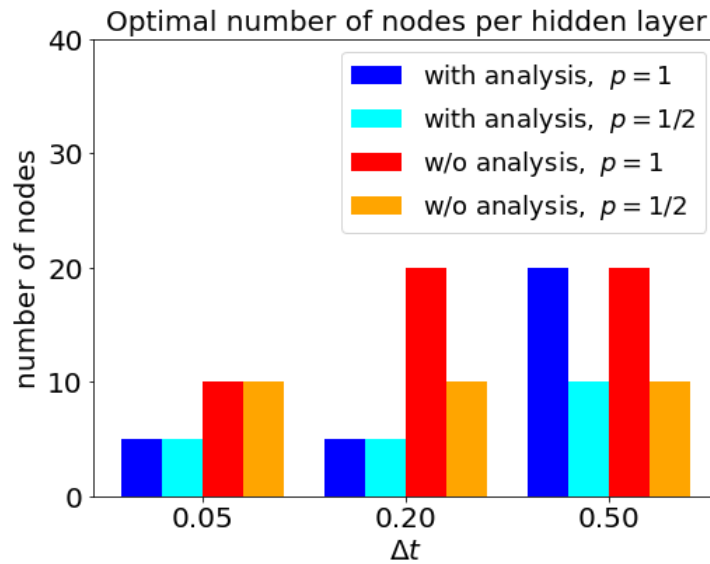
913 (b)



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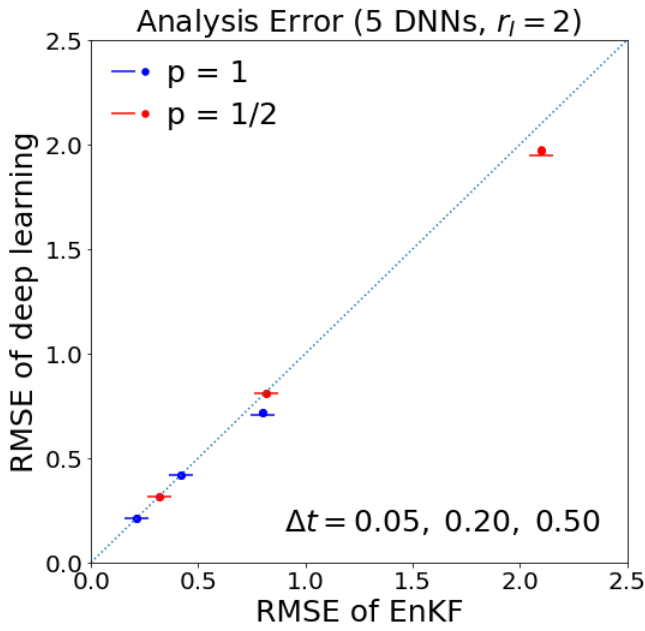
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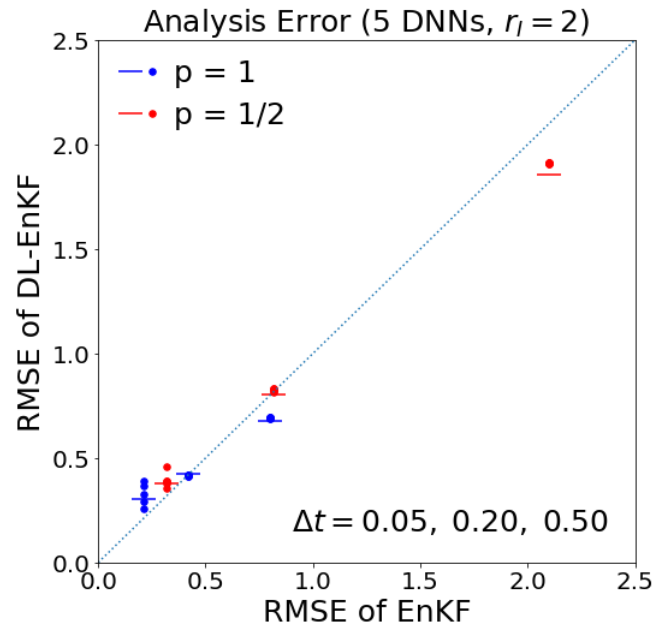
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(a)



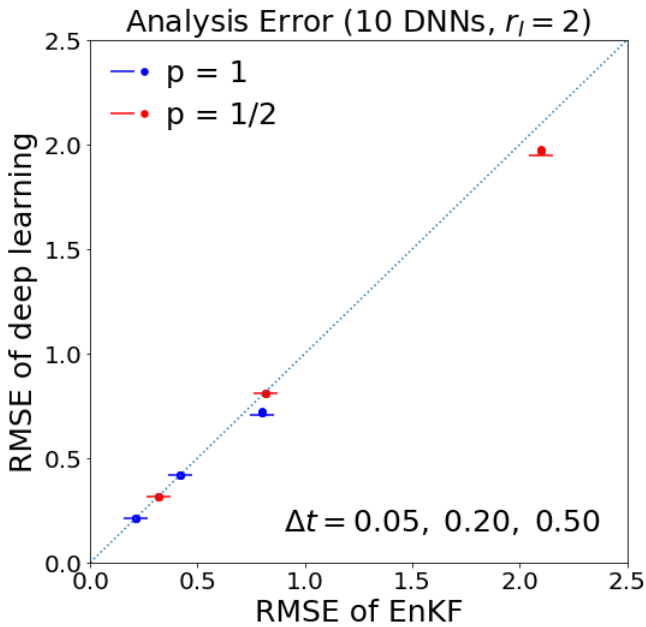
(b)



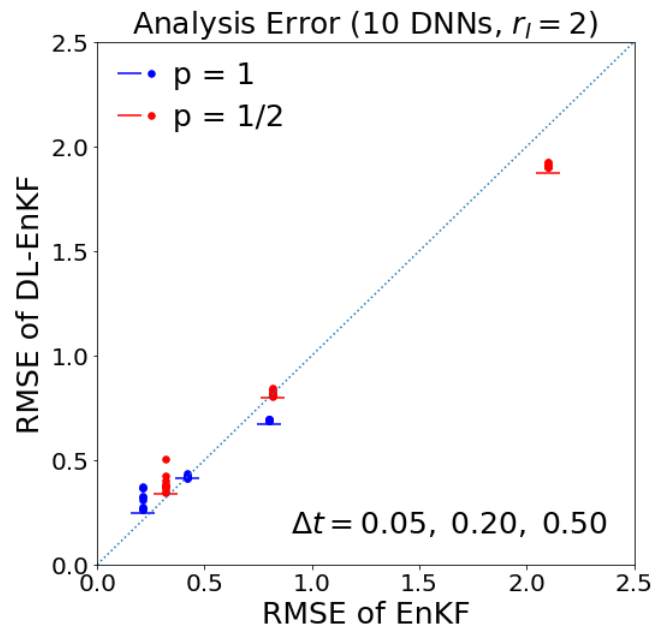
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(c)



(d)



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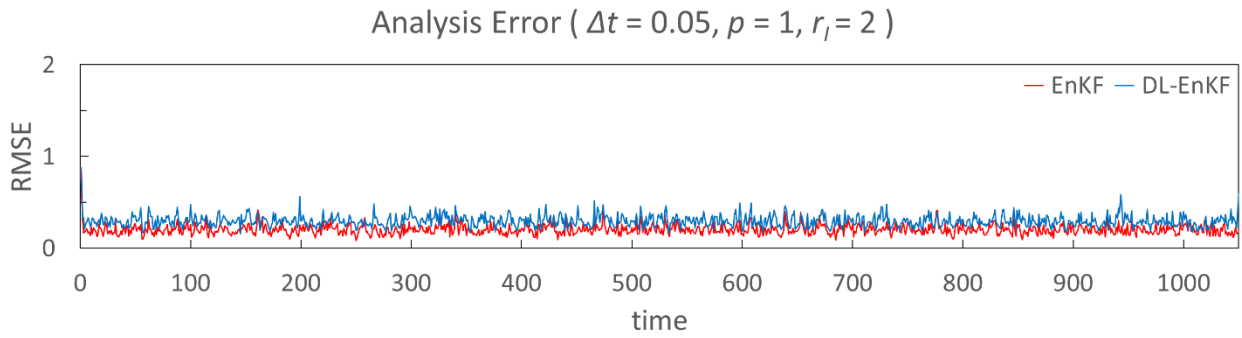
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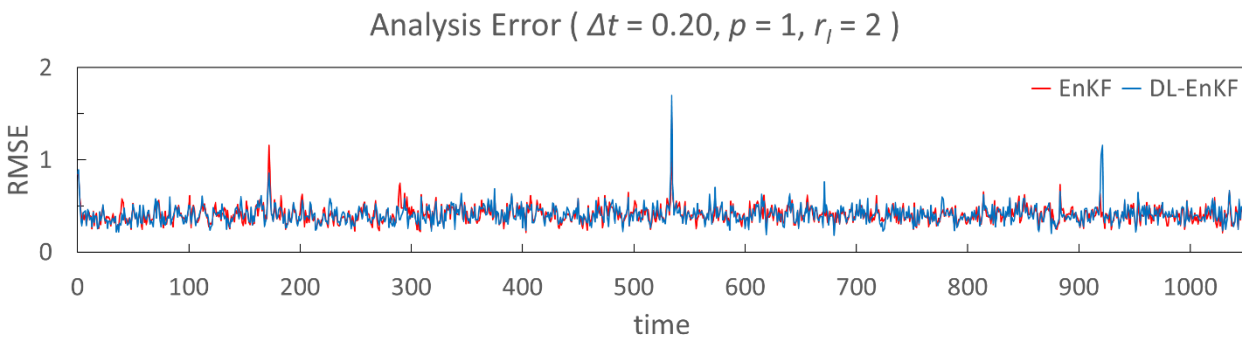
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(a)



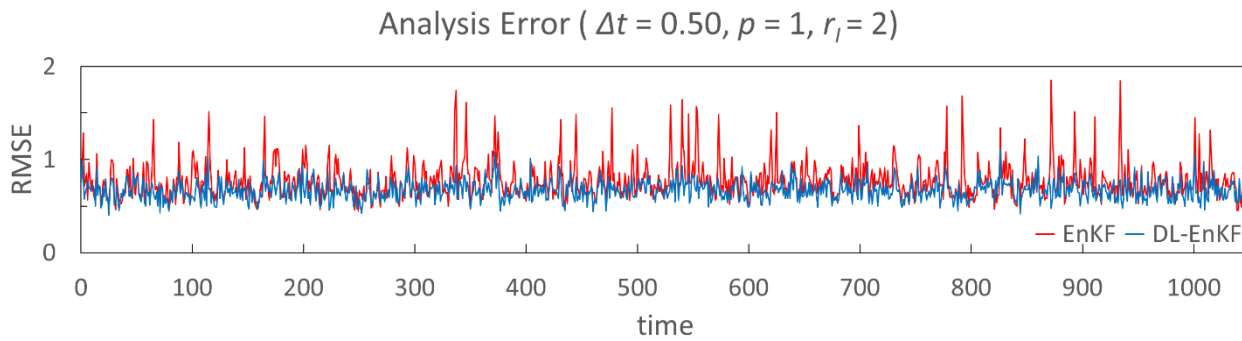
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(b)



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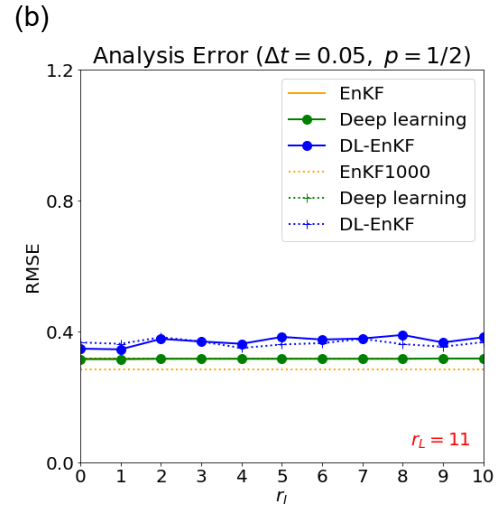
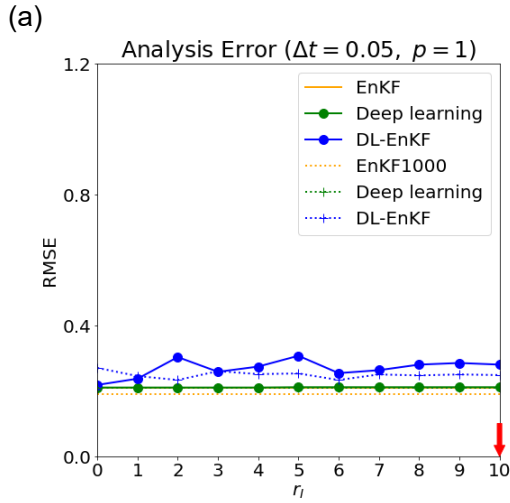


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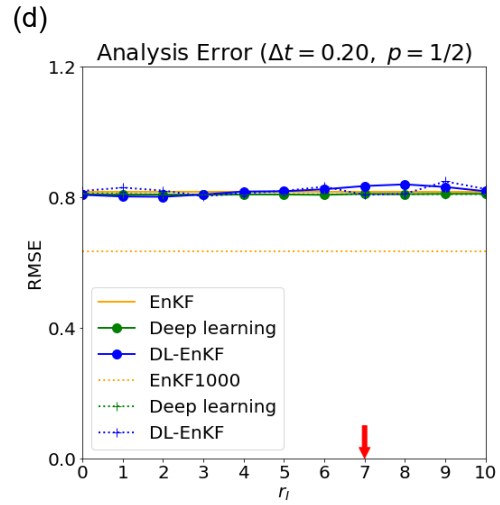
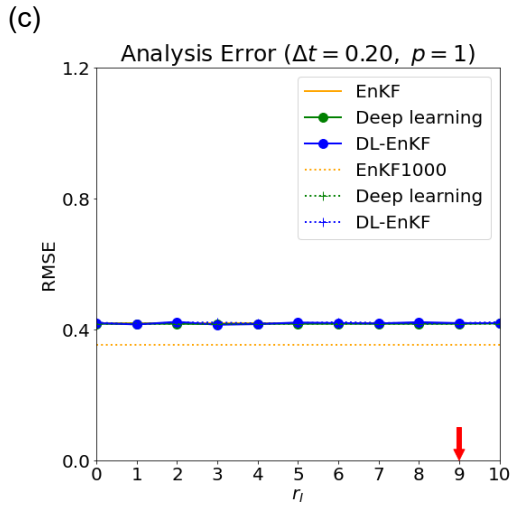
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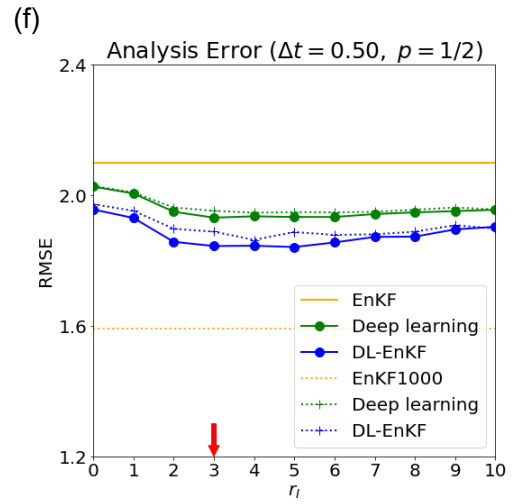
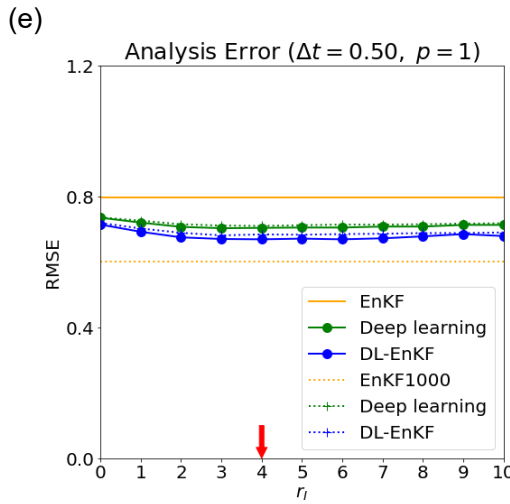
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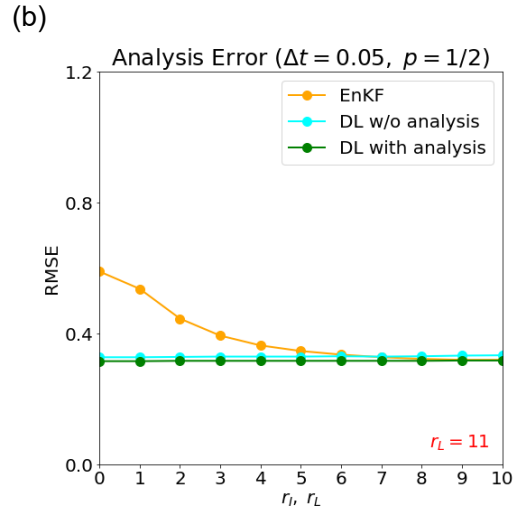
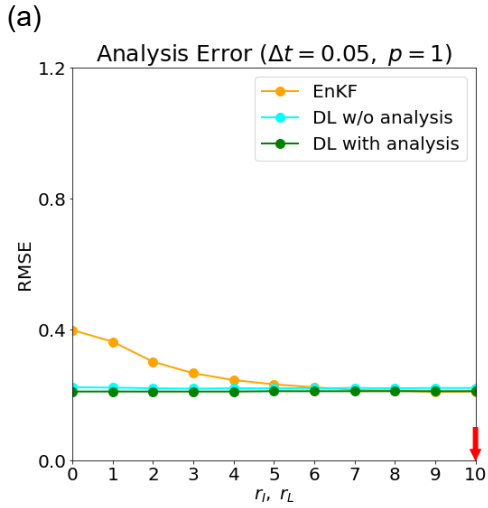
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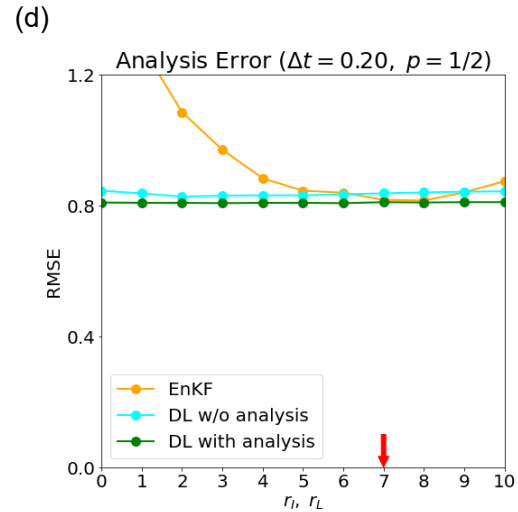
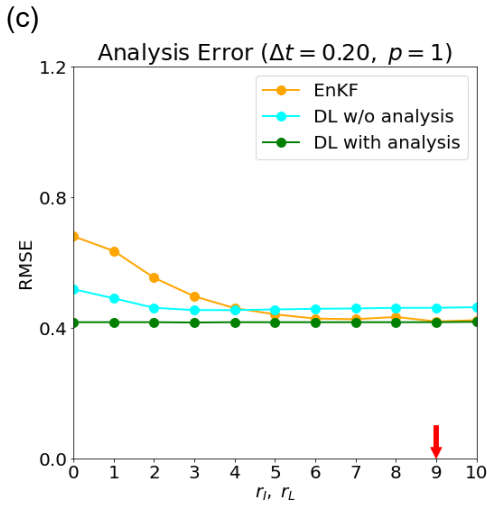
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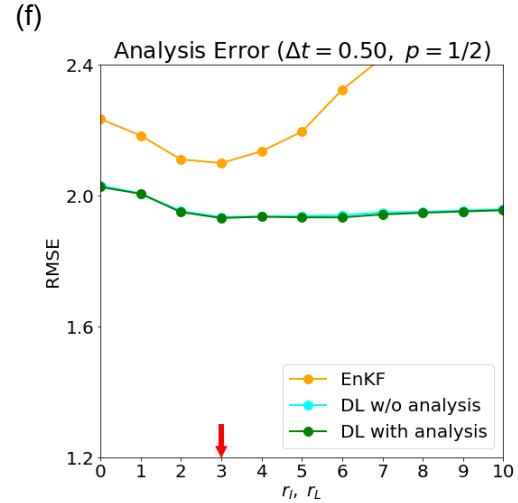
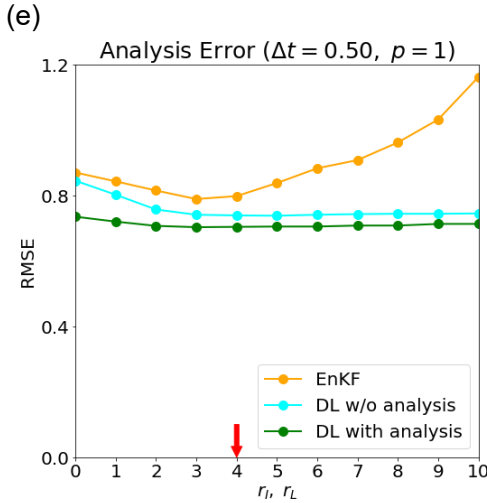
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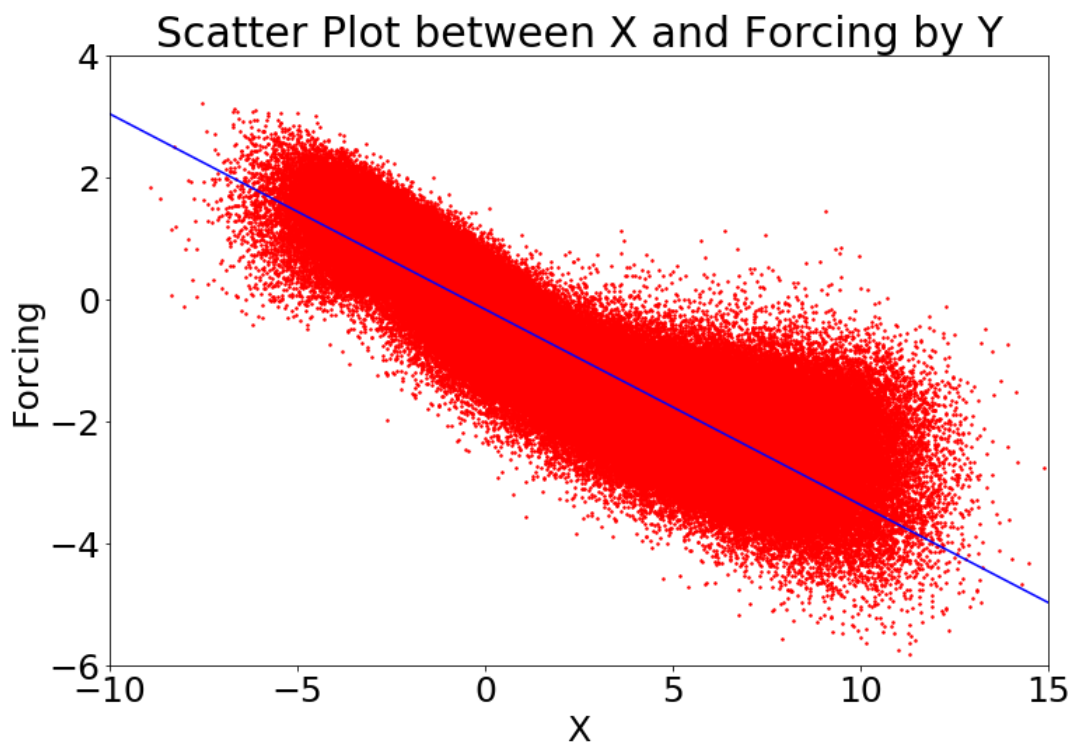


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962 Fig. 9 Comparison of RMSEs of EnKF (orange line), deep learning not including EnKF  
 963 analysis in input (cyan line), and deep learning including the EnKF analysis in input  
 964 (green line) for Exp-PA. The RMSE of EnKF is computed for each localization radius. The  
 965 abscissa is the input radius for deep learning and the localization radius for the EnKF. A  
 966 red arrow indicates the optimal localization radius. The observation time interval and the  
 967 probability of observations are (a) 0.05 and 1, (b) 0.05 and 1/2, (c) 0.20 and 1, (d) 0.20  
 968 and 1/2, (e) 0.50 and 1, (f) 0.50 and 1/2, respectively.

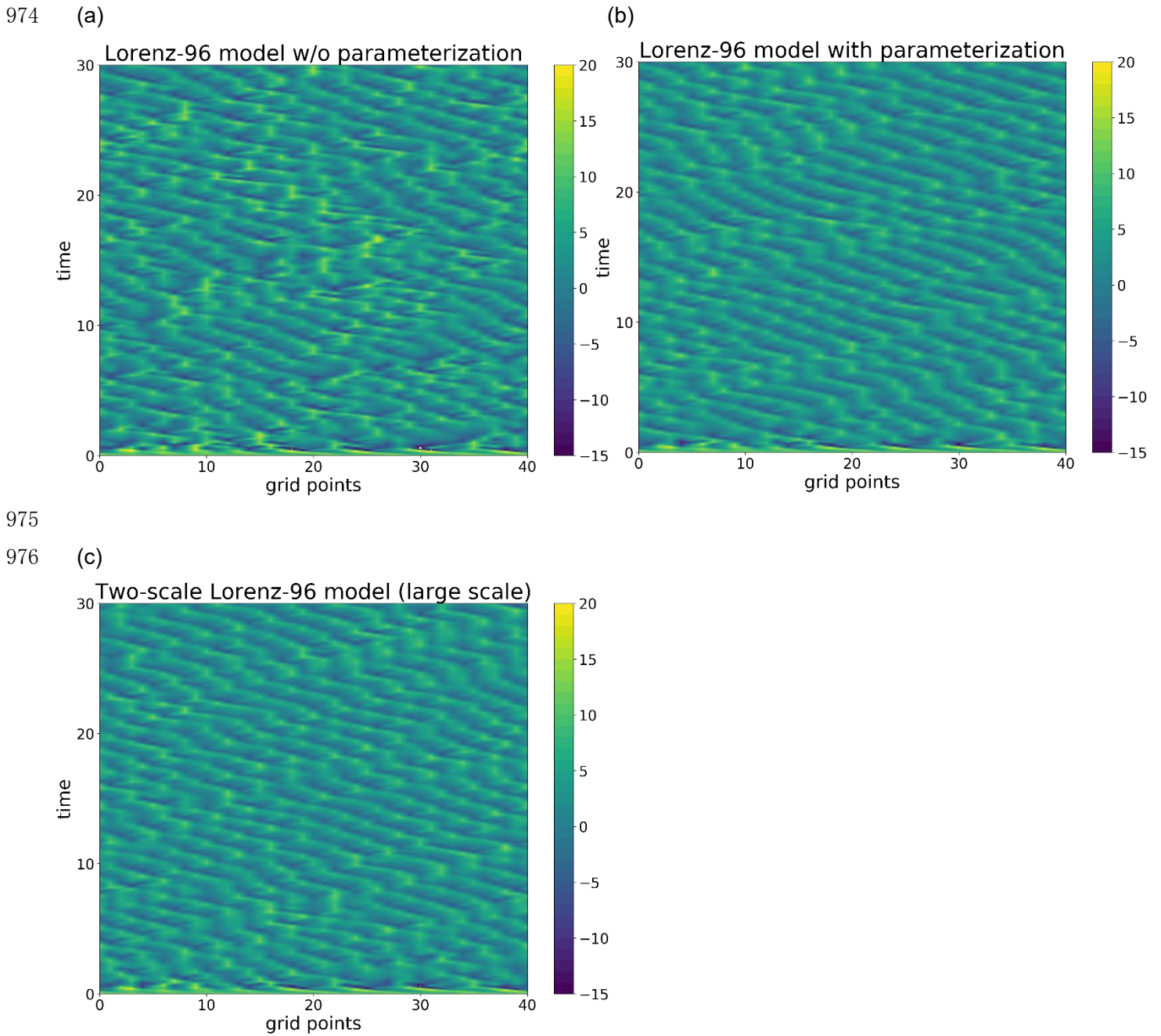




969

970 Fig. 10 Scatter plot between large-scale variables (abscissa) and large-scale forcing by  
971 small-scale variables (ordinate) of the two-scale Lorenz 96 model. A solid line is the result  
972 of linear function fitting.

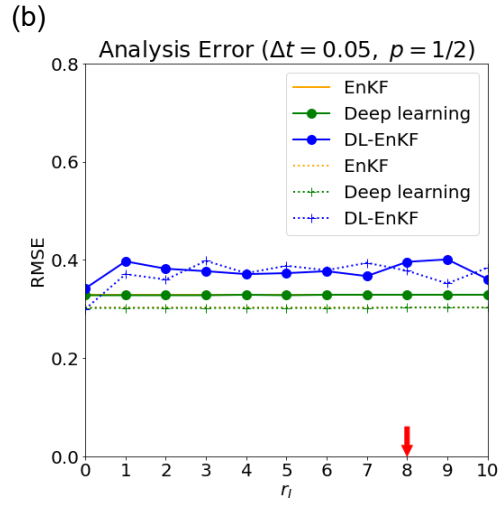
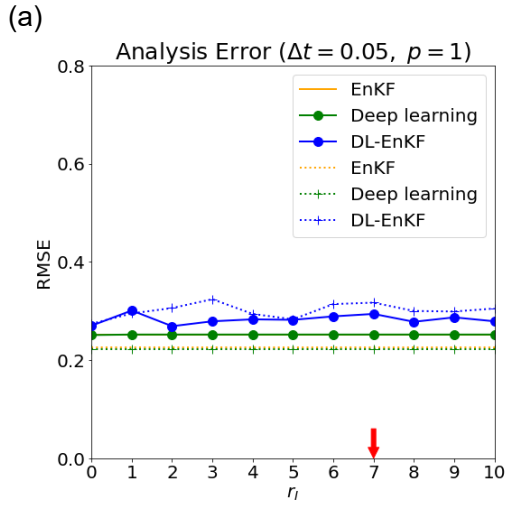
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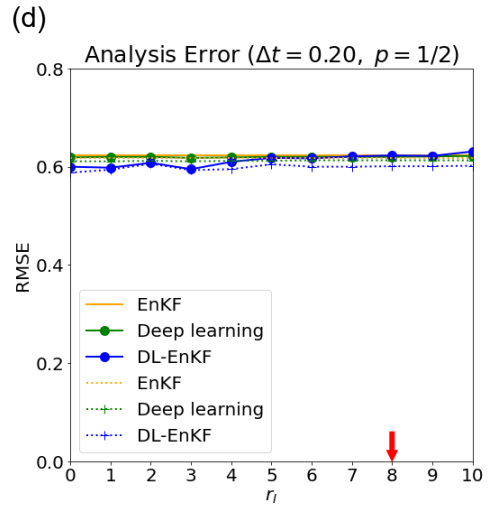
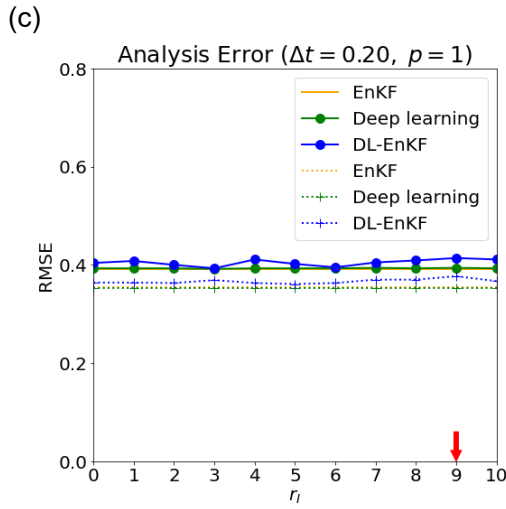
978 Fig. 11 Hovmöller diagrams of (a) the Lorenz 96 model, (b) the parametrized Lorenz 96  
 979 model, and (c) large-scale variables of the two-scale Lorenz 96 model.

980

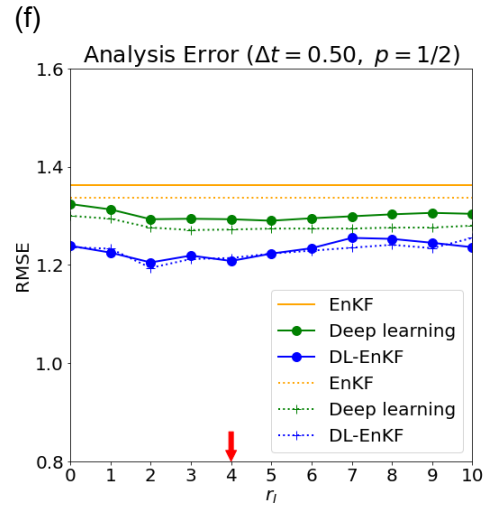
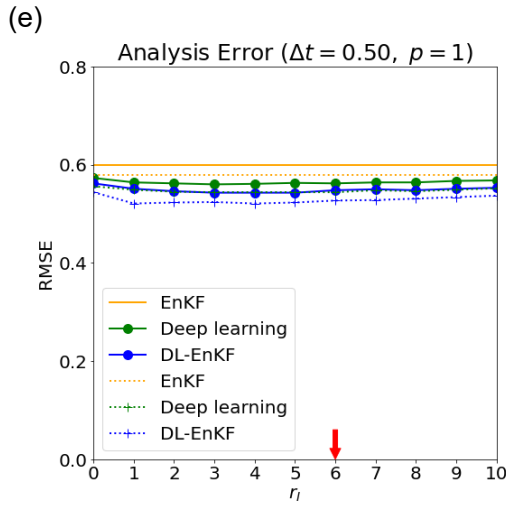
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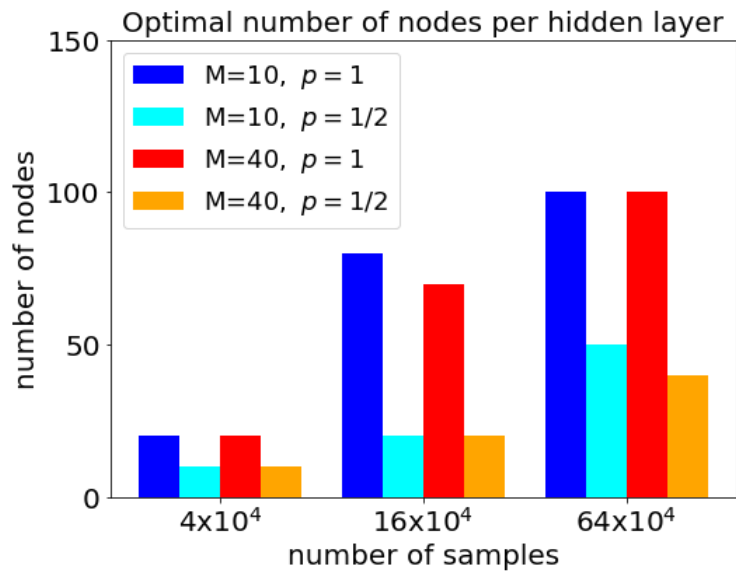


983



984

985 Fig. 12 Same as Fig. 8 except for Exp-IA (solid lines) and Exp-IB (broken lines) and that  
 986 an orange broken line indicates the RMSE of EnKF using the two-scale Lorenz 96 model.  
 987

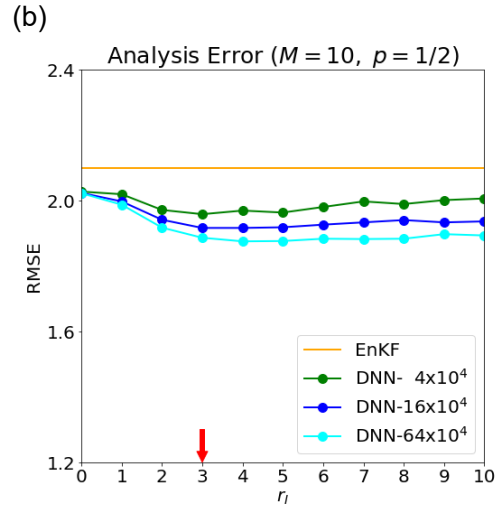
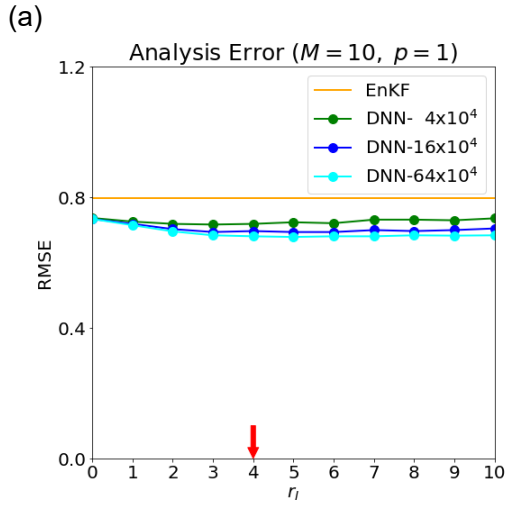


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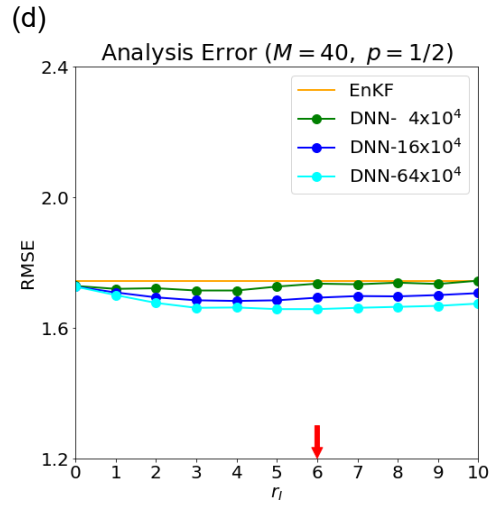
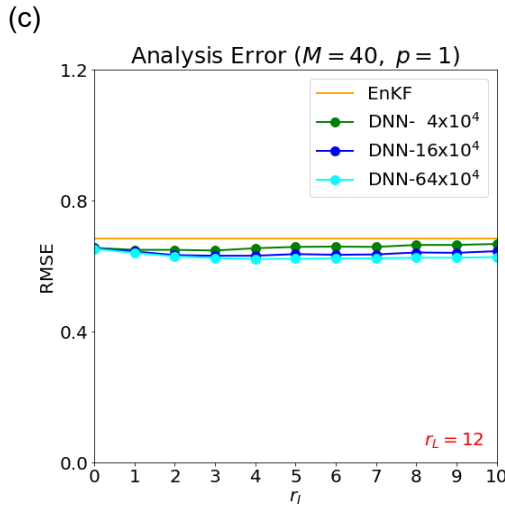
989 Fig. A1 The optimal number of nodes per hidden layer of a DNN with 5 hidden layers for  
 990 the time interval between observations of 0.50. The abscissa is the number of samples.  
 991 Blue and cyan bars are for the EnKF ensemble size of 10 for the probability of  
 992 observations 1 and 1/2, respectively. Red and orange bars are for the EnKF ensemble  
 993 size of 40 for the probability of observations 1 and 1/2, respectively.

994

995



996



997

998 Fig. A2 Comparison of RMSE between EnKF (orange lines) and the output of a DNN with  
 999 the number of samples of 40 000 (green lines), 160 000 (blue), and 640 000 (cyan) for  
 1000 the observation time interval of 0.50. The ensemble size of EnKF and the probability of  
 1001 observations are (a) 10 and 1, (b) 10 and 1/2, (c) 40 and 1, (d) 40 and 1/2, respectively.  
 1002 The RMSEs are plotted against the input radius, and a red arrow indicates the optimal  
 1003 localization radius of EnKF.

1004

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1010	

1011 Table 1 Models used in the training and test phases in the experiments.

1012

1013 (a) Perfect model experiments

	Exp-PA		Exp-PB	
	Training	Test	Training	Test
1016 Target	L	L	Analysis	L
1017 Observations	L	L	L	L
1018 Forecast ensemble	L	L	L	L

1019 L: Lorenz 96 model

1020 (b) Imperfect model experiments

	Exp-IA		Exp-IB	
	Training	Test	Training	Test
1023 Target	P	T	T	T
1024 Observations	P	T	T	T
1025 Forecast ensemble	P	P	T	T

1026 P: parameterized Lorenz 96 model, T: two-scale Lorenz 96 model

1027

1028 **Table 2** Architecture and training of the feedforward neural network.

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1030	No. of nodes of input layer	$3(2r_l + 1)$ for $p = 1$
1031		$4(2r_l + 1)$ for $p = 1/2$
1032	No. of hidden layers	5
1033	No. of nodes per hidden layer	5, 10, or 20 (optimized)
1034	No. of nodes of output layer	1
1035	Activation function	ReLU
1036	Loss function	Sum of squared error
1037	Gradient descent method	Adam*
1038	Learning rate	0.01 to 0.0001 (linear decay)
1039	No. of samples	40 000
1040	Mini-batch size	100
1041	No. of epochs	100

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\*: Kingma and Ba (2014)

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