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1	Geometry of rainfall ensemble means: from arithmetic
2	averages to Gaussian-Hellinger barycenters in
3	unbalanced optimal transport
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5	Le Duc <sup>1</sup> ,
6 7 8	Institute of Engineering Innovation, University of Tokyo Meteorological Research Institute, Tsukuba
9 10	and
10	Yohei Sawada
$\frac{12}{13}$	Institute of Engineering Innovation, University of Tokyo Meteorological Research Institute, Tsukuba
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<sup>&</sup>lt;sup>1</sup> Corresponding author address: Le Duc, Institute of Engineering Innovation, University of Tokyo, 1-1 Yaoi, Bunkyo-ku, Tokyo 305-0052.

E-mail: leduc@sogo.t.u-tokyo.ac.jp

#### Abstract

It is well-known in rainfall ensemble forecasts that ensemble means suffer substantially 22from the diffusion effect resulting from the averaging operator. Therefore, ensemble means 23are rarely used in practice. The use of the arithmetic average to compute ensemble means 24is equivalent to the definition of ensemble means as centers of mass or barycenters of all 25ensemble members where each ensemble member is considered as a point in a 26high-dimensional Euclidean space. This study uses the limitation of ensemble means as 27evidence to support the viewpoint that the geometry of rainfall distributions is not the 28familiar Euclidean space, but a different space. The rigorously mathematical theory 29underlying this space has already been developed in the theory of optimal transport (OT) 30 with various applications in data science. 31

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In the theory of OT, all distributions are required to have the same total mass. This 33 requirement is rarely satisfied in rainfall ensemble forecasts. We, therefore, develop the 34geometry of rainfall distributions from an extension of OT called unbalanced OT. This 35geometry is associated with the Gaussian-Hellinger (GH) distance, defined as the optimal 36cost to push a source distribution to a destination distribution with penalties on the mass 37discrepancy between mass transportation and original mass distributions. Applications of 38 the new geometry of rainfall distributions in practice are enabled by the fast and scalable 39 40 Sinkhorn-Knopp algorithms, in which GH distances or GH barycenters can be  $\mathbf{2}$ 

41	approximated in real-time. In the new geometry, ensemble means are identified with GH
42	barycenters, and the diffusion effect, as in the case of arithmetic means, is avoided. New
43	ensemble means being placed side-by-side with deterministic forecasts provide useful
44	information for forecasters in decision-making.

### 46 **1. Introduction**

Nowadays, ensemble forecasts play a vital role in numerical weather prediction. With 47advanced high-performance computing, the number of ensemble members is anticipated 48 to continue increasing in the future. Therefore, it is important to extract useful information 49from a large number of individual forecasts that give equally possible realizations. The 50standard technique in ensemble forecast is to reduce all ensemble members to a small 51number of important fields, such as quantile and probabilistic maps, ensemble means and 52spreads, and ensemble clusters. Quantile and probabilistic maps turn the high-dimensional 53probability distribution from an ensemble forecast into scalar quantities derived from 54one-dimensional marginal distributions at grid points. In contrast, ensemble means and 55ensemble clusters retain the high dimensionality of forecast fields and their coherent 56structures, but only show typical representatives of all ensemble members. This is the 57multivariate nature of the latter approach that makes the resulting forecast fields more 58interesting and complicated in use. At the same time, it opens rooms for new 59interpretations and explorations. In this study, we demonstrate this interesting problem with 60 the ensemble mean. 61

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The ensemble mean is chosen due to its simplicity, which is, in fact, the arithmetic average of all forecast members. This operator tends to filter out random noise but, at the same time, diffuse informative processes in individual members, leading to a smooth mean

66	field. The diffusion effect is noticeably clear when this operator is applied to rainfall. The
67	resulting rain field tends to spread out and is noticeably different from each member. Figure
68	1 illustrates this phenomenon with 20 rainfall forecasts over Kyushu Japan from the
69	mesoscale ensemble prediction system MEPS of Japan Meteorology Agency (JMA) (Ono
70	et al., 2021) using the JMA's operational limited area model ASUCA (Ishida et al., 2022).
71	The arithmetic mean of 20 rainfall distributions differs greatly from the deterministic
72	forecast. As a result, ensemble means of rainfall are rarely used in practice.

Figure 2a conceptually explains this fact using two one-dimensional rainfall distributions 74with the same shapes but a small displacement error. We expect that the mean should 75retain a similar shape with its location between the locations of the two individual 76distributions. However, the arithmetic average yields an undesirable result: a bimodal 77distribution with peaks much smaller than those of the two members. Furthermore, the 78mean distribution covers a wide area spreading from the left leg of the first member to the 79right leg of the second member. Although this explanation is employed in a 80 one-dimensional space, it can be carried out in higher spaces without any significant 81 difference. 82

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84 An important reason for using the ensemble mean as a representative forecast is that if 85 the probability distribution of forecasts is a multivariate normal distribution, the forecast

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mean also gives the mode of this probability distribution. As a result, the forecast mean 86 becomes the most probable forecast and can be taken as the best approximation of the 87 true state. Under this assumption, the expected error between the ensemble mean and the 88 observation is proved to be proportional to the reciprocal of the square root of the number 89 of ensemble members (see Appendix A). Thus, by increasing the number of ensemble 90 members, we expect to obtain a more accurate forecast through the ensemble mean. 91However, this is not the case even with a relatively large number of ensemble members, as 92observed in Fig. 3a (1000 members in this case). Instead of two-dimensional rainfall 93distributions, the problem is simplified by only plotting the time series of 1-hour precipitation 94averaged over the Ichifusa catchment in Kyushu Japan. The forecasts are obtained from a 951000-member ensemble prediction system LETKF1000 (Duc et al., 2021) using the JMA's 96 former operational limited area model NHM (Saito et al., 2006). Like the two-dimensional 97case, the ensemble mean does not show a similar pattern as the corresponding time series 98 from the observations. However, intriguingly, if we plot the time series of accumulated 99 rainfall instead of rain rates (Fig. 3b), the ensemble mean becomes nearly identical to the 100 101accumulated rainfall observations. This striking fact has been observed in several studies using a large number of ensemble members (Kobayashi et al. 2020, 2023) without any 102adequate explanation. 103

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Geometrically, if we consider each distribution with *n* elements as a point in an

n-dimensional space  $\mathbb{R}^n$ , the ensemble mean is simply the center of mass or the 106 barycenter of all members, assuming that all members have the same mass. Figure 3b 107108implies that if we want to retain the meaning of the ensemble mean as a barycenter, we need to work in the space of cumulative rainfall distributions. In other words, the 109appropriate use of the ensemble mean of rainfall in one-dimensional cases is with 110accumulated rainfall distributions rather than rain rate distributions. We can easily verify the 111 validity of this hypothesis with the simple example in Fig. 2a. Figure 2b plots the cumulative 112distributions corresponding to the distributions in Fig. 2a. As expected, the cumulative 113distribution of the expected mean lies between the two cumulative distribution members. 114However, what is notable here is that in cumulative forms, the expected mean is nearly 115identical to the arithmetic mean of the inverses of the cumulative distributions. Note that in 116Fig. 2b, since the cumulative distributions are one-to-one maps, their graphs also represent 117their inverses, in which time is considered as a function of cumulative rainfall. 118

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Figure 2b suggests the following procedure to find the expected ensemble mean for anynumber of one-dimensional distributions:

122 (1) Transforming all distributions to the space of inverses of cumulative distributions;

123 (2) Calculating the arithmetic mean in the transformed space;

124 (3) Transforming the resulting ensemble mean back to the space of distributions.

125 Of course, the important problem is how we can justify such a procedure with a robust

theoretical base. Note that Fig. 2b demonstrates the first two steps (1) and (2) of this
 procedure, while Fig. 2a demonstrates the last step (3).

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The above procedure shows a potential way of finding expected ensemble means in 129high-dimensional cases. Thus, all we need is finding an appropriate space in which the 130ensemble mean retains its meaning as barycenters while being robust to the diffusion 131effect in the presence of displacement errors among members. However, an attempt to 132extend cumulative distributions from one-dimensional cases to high-dimensional cases 133does not work since it is unclear how to define a high-dimensional cumulative field from a 134high-dimensional distribution. A natural question is whether this space exists at all in 135136 general cases. It is noted that even in one-dimensional cases, the arithmetic average operator only makes sense if all one-dimensional distributions have the same total mass. 137This means that even in the simplest cases, the existence of such a space is still 138questionable. 139

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In this study, we show that the theory of optimal transport (OT) (Villani, 2009; Santambrogio, 2015; Peyré and Cuturi, 2019) proposes an elegant solution to this problem. Instead of working in a transformed space, we continue to stick with the space of rainfall distributions but endowed with a new geometry. Thus, the similarity between any distributions is no longer measured by the normal Euclidean distance but is replaced by a

new distance defined in the context of OT. In particular, this distance becomes the normal
Euclidean distance between inverses of cumulative distributions in one-dimensional cases.
The theory of OT relevant to this study, i.e., unbalanced OT, is presented in the next
section. The barycenters, resulting from the new geometry of rainfall distributions defined
by unbalanced OT, are described and analyzed in Section 3. Finally, Section 4
summarizes the main points of this study and discusses some potential applications.

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#### **153 2. Unbalanced optimal transport**

Although OT has been successfully applied in many fields of data science, there is only a 154limited number of applications of OT in geosciences (Farchi et al., 2016; Métivier et al., 1552016; Yang et al., 2018; Sambridge et al., 2022). However, it is worth noting that in recent 156years, we have seen an increase in studies using OT in geophysical data assimilation 157(Reich and Cotter, 2015; Feyeux et al., 2018; Li et al., 2018; Tamang et al., 2021; 158Vanderbecken et al., 2023). In order to make OT accessible to the meteorology community, 159the theory of OT will be adapted to rainfall distributions and simplified in this section. A 160more rigorous treatment for probabilistic distributions can be found in Villani (2009) and 161Santambrogio (2015). Our mathematical treatment in this section mainly follows Peyré and 162Cuturi (2019). 163

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165 Let two vectors  $\mathbf{a}, \mathbf{b} \in \mathbb{R}^n_+$  denote two rainfall distributions with the same total rain mass

over the same domain D. We consider rainfall in the same domain, however, in the theory 166 of OT, two distributions need not be on the same domain. We call a matrix  $\mathbf{P} \in \mathbb{R}^{n \times n}_+$  a 167transport plan that moves **a** to **b** in the sense that the element  $P_{ij}$  denotes the rain mass 168from the bin *i* to the bin *j*. A bin corresponds to a grid box in the domain. For mass 169conservation, we impose two constraints on the elements of P 170P1 = a, (1a) 171 $\mathbf{P}^{\mathrm{T}}\mathbf{1}=\mathbf{b},$ 172(1b) where 1 denotes a vector with all elements equal one. From the constraints (1), it is easy 173to verify that the total rain mass is conserved 174 $\mathbf{1}^{\mathrm{T}}\mathbf{a} = \mathbf{1}^{\mathrm{T}}\mathbf{P}\mathbf{1} = (\mathbf{P}\mathbf{1})^{\mathrm{T}}\mathbf{1} = \mathbf{1}^{\mathrm{T}}\mathbf{P}^{\mathrm{T}}\mathbf{1} = \mathbf{1}^{\mathrm{T}}\mathbf{b}.$ (2)175Associated with **P**, we have a matrix  $\mathbf{C} \in \mathbb{R}^{n \times n}_+$  whose element  $C_{ij}$  denotes the 176177transportation cost from the bin *i* to the bin *j*. 178The original theory of OT seeks the OT plan  $P^*$  that minimizes the following objective 179function 180  $L_{\mathbf{C}}(\mathbf{a},\mathbf{b}) = \min_{\mathbf{P}} \langle \mathbf{C},\mathbf{P} \rangle = \min_{\mathbf{P}} \sum C_{ij} P_{ij},$ (3)181 subject to the constraints (1) where the symbol () denotes the inner product of two 182

matrices. With the linear constraints (1), the linear programming problem (3) is convex and therefore has a global minimum. This formulation is known as the Kantorovich problem in OT (Kantorovich, 1942), an extension of the Monge problem in which mass splitting is not

186	allowed (the entire mass from a bin is moved to another bin). What is the connection
187	between the optimal cost $L_{C}(a,b)$ and the geometry of rainfall distributions?
188	
189	The connection follows from an important theorem: if $C_{ij} = \ \mathbf{x}_i - \mathbf{x}_j\ _p^p$ where $\ \mathbf{x}_i - \mathbf{x}_j\ _p =$
190	$D_{ij}$ is the Minkowski distance of order p between two grid points $\mathbf{x}_i, \mathbf{x}_j$ , then $L_{\mathbf{C}}(\mathbf{a}, \mathbf{b})$ induces
191	a distance between $\mathbf{a}$ and $\mathbf{b}$ , which is called the p-Wasserstein distance
192	$W_p(\mathbf{a},\mathbf{b}) = L_{\mathbf{D}^p}(\mathbf{a},\mathbf{b})^{1/p}.$ (4)
193	We replace <b>C</b> with $\mathbf{D}^p$ in (4) to emphasize that the new distance is defined by using the
194	Minkowski distance as the transportation cost. Also, recall that the Minkowski distance of
195	order 2 is the familiar Euclidean distance. In order to be a distance, the distance function
196	has to be positive, symmetric, and has to obey the triangular inequality. The p-Wasserstein
197	distance usually does not have an analytic form, and can only be estimated numerically.
198	
199	However, for one-dimensional distributions, $W_p(\mathbf{a}, \mathbf{b})$ has a closed form given by
200	$W_p(\mathbf{a},\mathbf{b}) = \ F^{-1}(\mathbf{a}) - F^{-1}(\mathbf{b})\ _p$ , (5)
201	where $F^{-1}(\mathbf{a})$ denotes the inverse of the cumulative distribution constructed from $\mathbf{a}$ . Thus,
202	when $p = 2$ , i.e., $\ \mathbf{x}_i - \mathbf{x}_j\ _2$ is the Euclidean distance between $\mathbf{x}_i$ and $\mathbf{x}_j$ , the
203	2-Wasserstein distance is identical to the Euclidean distance between $F^{-1}(\mathbf{a})$ and $F^{-1}(\mathbf{b})$ .
204	This fact supports our averaging operation in Fig. 2b, where we take the arithmetic mean of
205	two inverses of cumulative distributions. This step is equivalent to taking the barycenter ${f c}$

of two distributions **a** and **b** with respect to the 2-Wasserstein distance

207 
$$\min_{\mathbf{c}} \left[ 0.5 W_2(\mathbf{a}, \mathbf{c})^2 + 0.5 W_2(\mathbf{c}, \mathbf{b})^2 \right].$$
 (6)

Of course, in practice, we avoid minimizing (6) by simply taking the average of  $F^{-1}(\mathbf{a})$ and  $F^{-1}(\mathbf{b})$ , and we transform it back to the space of distributions.

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Recall that the OT problem (3) strictly requires the same total mass between **a** and **b**. However, rainfall distributions from different ensemble members rarely satisfy this condition. Therefore, rainfall distributions cannot be considered in the same Wasserstein space. This limitation prevents the application of OT for rainfall distributions. Clearly, the averaging operation does not make sense in Fig. 2b if the two cumulative distributions have two different heights.

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There are many attempts to relax the requirement of the same total mass in OT. The most successful approach is known under the name unbalanced optimal transport (UOT) (Frogner et al., 2015; Chizat et al., 2018a; Liero et al., 2018), which relaxes the objective function (3) with the new form

222 
$$L_{\mathbf{C}}(\mathbf{a},\mathbf{b}) = \min_{\mathbf{P}} \left[ \langle \mathbf{C}, \mathbf{P} \rangle + \tau \mathrm{KL}(\mathbf{P}\mathbf{1},\mathbf{a}) + \tau \mathrm{KL}(\mathbf{P}^{\mathrm{T}}\mathbf{1},\mathbf{b}) \right],$$
 (7)

where  $\tau$  is the marginal relaxation parameter that penalizes the discrepancy between mass transportation and original mass distributions. This discrepancy is measured with the

226 
$$\operatorname{KL}(\mathbf{a},\mathbf{b}) = \sum_{i} [a_{i} \log (a_{i}/b_{i}) + b_{i} - a_{i}].$$
(8)

The existence of the term  $b_i - a_i$  in (8) is due to unequal total mass between **a** and **b**. Thus, (7) replaces the strong constraints (1) with weak constraints through the marginal relaxation terms, and controls these constraints by  $\tau$ . It can be verified that (7) reduces to (3) in the limit  $\tau \rightarrow \infty$  under the total mass constraint  $\mathbf{1}^{\mathrm{T}}\mathbf{a} = \mathbf{1}^{\mathrm{T}}\mathbf{b}$ .

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Like the 2-Wasserstein distance, when  $C_{ij} = ||\mathbf{x}_i - \mathbf{x}_j||_2^2$  is the square of the Euclidean distance, the optimal cost (7) defines a distance between **a** and **b**, called the Gaussian-Hellinger (GH) distance

235 GH(**a**,**b**) = 
$$L_{\mathbf{D}^2}(\mathbf{a},\mathbf{b})^{1/2}$$
. (9)

However, unlike  $W_p(\mathbf{a}, \mathbf{b})$ ,  $GH(\mathbf{a}, \mathbf{b})$  can only be estimated numerically even for one-dimensional distributions. Using this new distance, we can estimate the barycenter of an ensemble of rainfall distributions  $\{\mathbf{a}_k\}_{k=1,K}$  from the following minimization problem  $\min_{\mathbf{b}} \left[ \sum_{k,K} \frac{1}{K} GH(\mathbf{a}_k, \mathbf{b})^2 \right]$ . (10)

In other words, we seek the distribution **b** that minimizes the averaged GH distances from **b** to all  $\mathbf{a}_k$ . Figure 4 illustrates this barycenter problem by showing GH-barycenters of two different rainfall distributions in terms of both volumes and spreads. The computation is employed by using the algorithm described in the next section.

**3. Regularized Gaussian-Hellinger barycenters** 

The most common method to solve the linear programing problem (3) is the simplex 246algorithm. In general, the computational complexity of the simplex method is  $O(n^3 logn)$ 247which limits the applications of OT in practice. Recall that n denotes the size of the vectors 248a,b, which is the number of grid points in the domain D. With the introduction of the 249marginal relaxation, the UOT minimization problem (7) is no longer a linear programing 250problem. The closed solution exists when distributions have Gaussian forms (Janati et al., 2512020). Blondel et al. (2018) proposed to solve this using the L-BFGS-B algorithm with the 252squared norm in place of the KL divergence. Sato et al. (2020) provided an effective 253solution when distributions have a tree structure. Chapel et al. (2021) showed that (7) can 254be turned into a non-negative linear regression problem, and solved with non-negative 255matrix factorization. However, what makes UOT applicable in practice is the introduction of 256entropic regularization into UOT, leading to scalable and fast algorithms. 257

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The idea of adding an entropic regularization term into (3) has been proposed by Cuturi (2012) to make the problem strictly convex and, therefore, simpler to solve. In particular, regularization enables the use of the fast Sinkhorn-Knopp algorithm (Sinkhorn and Knopp, 1967; Benamou et al., 2015) with the complexity  $O(n^2)$  to approximate the optimal plan. This idea has been introduced into UOT by Chizat et al. (2018b) to derive matrix scaling algorithms for UOT problems in the vein of the Sinkhorn-Knopp algorithm. Thus, we

regularize the minimization problem (7) by adding an entropic regularization term  
L{
$$(a,b) = \min_{\mathbf{P}} [(\mathbf{C},\mathbf{P}) + \tau KL(\mathbf{P1},a) + \tau KL(\mathbf{P^T1},b) - \varepsilon H(\mathbf{P})],$$
 (11)  
where  $\varepsilon$  is the entropic regularization parameter, and H(P) represents the entropy of P  
H(P) =  $-\sum_{ij} [P_{ij} \log (P_{ij}) - P_{ij}].$  (12)  
The regularized GH distance associated with the regularized UOT problem (11) is  
expressed as GH <sup>$\varepsilon$</sup> (a,b), which transforms our barycenter problem (10) to  
min  $\left[\sum_{k\bar{k}} \frac{1}{6} GH^{\varepsilon}(a,k,b)^{2}\right].$  (13)  
By letting  $\varepsilon$  go to zero in minimizing (13), we obtain the GH barycenter of the original  
barycenter problem (10).  
The matrix scaling algorithm developed by Chizat et al. (2018b) for the barycenter  
problem (13) is reproduced here with some adaptations  
Algorithm 1:  
Input:  $\varepsilon$ ,  $\tau$ ,  $\{a_k\}_{k=1,K}$ , C  
Local variables:  $\phi$ , K,  $\{u_k\}_{k=1,K}, \{v_k\}_{k=1,K}$   
Initialization:  $\phi = \tau/(\tau + \varepsilon)$ , K = exp ( $-C/\varepsilon$ ),  $\mathbf{v}_k \leftarrow \mathbf{1}$  ( $k = 1,K$ )

282 
$$\mathbf{u}_k \leftarrow [\mathbf{a}_k \oslash \mathbf{K} \mathbf{v}_k]^{\phi} \ (k = 1, K),$$
 (14a)

283 
$$\mathbf{b} \leftarrow \left[ \sum_{k\overline{k}} \frac{1}{(\mathbf{K}^{\mathrm{T}} \mathbf{u}_{k})^{1-\phi}} \right]^{1/(1-\phi)}, \qquad (14b)$$

284 
$$\mathbf{v}_k \leftarrow [\mathbf{b} \oslash \mathbf{K}^{\mathrm{T}} \mathbf{u}_k]^{\phi} (k = 1, K),$$
 (14c)

Until convergence

286 Output: **b** 

The symbol  $\oslash$  in (14a) and (14c) denotes the element-wise division operator. The matrix **K** is the element-wise exponential matrix of  $-C/\varepsilon$ . It is worth noting that this algorithm is prone to numerical underflow and overflow when  $\varepsilon$  is small. Therefore, it is better to work with  $-C/\varepsilon$ ,  $\log u_k$ ,  $\log v_k$ , than **K**,  $u_k$ ,  $v_k$ . This strategy is known as the log-domain Sinkhorn-Knopp algorithm.

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Figure 4 shows the impact of the entropic regularization and the marginal relaxation on 293approximations of GH barycenters using Algorithm 1. Due to the fact that OT plans become 294less sparse under the entropic regularization,  $\varepsilon$  with large values cause the mass spread 295in approximated GH barycenters, as illustrated in Fig. 4a. This implies that to avoid the 296diffusion effect,  $\varepsilon$  should be set to small values. However, if  $\varepsilon$  are too small, the 297Sinkhorn-Knopp algorithm will suffer from numerical underflow and overflow, and fail to 298terminate. For the marginal relaxation  $\tau$ , it is interesting to see that when a large mass 299300 discrepancy is enabled, i.e.,  $\tau$  are small, the UOT barycenter problem leads back to a distribution similar to the arithmetic mean of two ensemble members. Therefore,  $\tau$  should 301be set to large values to avoid arithmetic means. However, if  $\tau$  are too large, the algorithm 302will converge slower due to  $\phi \sim 1$ . The settings of  $\varepsilon = 10^{-4}$  and  $\tau = 10$  will be applied in 303 304 all the computations using Algorithm 1.

When applying the matrix scaling algorithm to two-dimensional rainfall distributions, the 306 307 most challenging issue is the exponential increase in computational costs. For the one-dimensional case in Fig. 4, the size of the problem is n = 100. Let us consider 308 two-dimensional distributions with the same size in each direction. The problem size 309 becomes  $n = 100^2 = 10^4$ , leading to an increase of  $10^8/10^4 = 10^4$  times in the 310 computational cost. Notice that the computation cost in Algorithm 1 is mainly dominated by 311the matrix-vector products,  $\mathbf{K}\mathbf{v}_k$  and  $\mathbf{K}^{\mathrm{T}}\mathbf{u}_k$ , which takes  $\mathcal{O}(n^2)$  operations. In order to 312mitigate the huge computational cost in two-dimensional cases, we model K as a tensor 313product  $\mathbf{K} = \mathbf{K}_x \otimes \mathbf{K}_v$ , so that the matrix-vector products can be done in each x, y direction 314separately. As a result, the computation cost reduces to  $O(n^{3/2})$  resulting in  $10^2$  time 315increase in the computational cost as compared to one-dimensional cases, which is 316 affordable in practice. 317

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Figure 5 shows the GH barycenters for consecutive 3-hour precipitation from 00 to 09 JST on July 4<sup>th</sup> 2020, estimated from 20 members of MEPS. Corresponding deterministic forecasts and arithmetic means are also plotted for comparison. Clearly, all GH barycenters avoid the diffusion effect as observed in the arithmetic means. Furthermore, the GH barycenters provide additionally useful information for the deterministic forecasts by showing locations where all members disagree with the deterministic forecasts,

therefore, overconfidence from the deterministic forecasts may be avoided. This information is important for forecasters in decision-making.

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Objective verification using the Fractions Skill Score (FSS) (Roberts and Lean, 2008) is 328 performed to quantify the performances of the two kinds of barycenters in addition to the 329 deterministic forecasts. The verification results are shown in Fig. 6. Clearly, for the 330arithmetic means, due to the diffusion effect, their FSSs drop rapidly to zero when the 331rainfall thresholds increase. This means the GH barycenters outperform the arithmetic 332means for intense rain. However, the arithmetic means are not entirely worse than the GH 333 barycenters. At the rainfall thresholds smaller than 10 mm (3h)<sup>-1</sup>, the arithmetic means 334yield forecasts slightly better than the GH barycenters. The reason can be traced back to 335 Fig. 1, where a large rainfall area forecasted by the arithmetic means (Fig. 1c) because the 336diffusion effect is unexpectedly in accordance with the observed rainfall area (Fig. 1a). In 337contrast, individual forecasts similar to the deterministic forecasts tend to predict a rainfall 338 area much smaller than that of the observations. As a result, the GH barycenters become 339 slightly worse than the arithmetic means in predicting the rainfall area. 340

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The computing program for the barycenters in Fig. 5 is parallelized along the direction of ensemble members, i.e., each processor only works with a subset of ensemble members. With 20 Intel Xeon processors and the domain consisting of 311x242 grid points, each GH

barycenter takes three minutes to calculate. The running time can be accelerated considerably if parallelization in the x and y directions is employed. Since all GH barycenters are independent with respect to different lead times, they can be produced in parallel. This enables GH barycenters to be deployed in real-time.

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#### **4.** Discussions and conclusions

It is well-known in rainfall ensemble forecasts that ensemble means suffer substantially 351from the diffusion effect resulting from the averaging operator. Therefore, ensemble means 352are usually not comparable with any ensemble members, and as a result, are rarely used 353in practice. The use of the arithmetic average to compute ensemble means is equivalent to 354355the definition of ensemble means as centers of mass of all ensemble members where each member is considered as a point in a high-dimensional Euclidean space. This study uses 356the limitation of ensemble means as evidence to support the viewpoint that the geometry of 357rainfall distributions is not the familiar Euclidean space, but a different metric space 358associated with a certain distance. The rigorously mathematical theory underlying this 359space has already been developed in the theory of OT with various applications in other 360 disciplines, of which objects are the same kind of distributions as rainfall distributions. 361

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In the theory of OT, all distributions are required to have the same total mass. This
 requirement is, of course, rarely satisfied in rainfall ensemble forecasts. We, therefore,

develop the geometry of rainfall distributions from an extension of OT called UOT. This 365 geometry is associated with the GH distance defined in UOT. This distance is the optimal 366 cost to push a source distribution to a destination distribution with penalties on the mass 367discrepancy between mass transportation and original mass distributions. The applications 368 of the new geometry of rainfall distributions in practice are enabled by the fast and scalable 369 Sinkhorn-Knopp algorithms, in which GH distances or GH barycenters can be 370approximated in real-time. By replacing arithmetic means with GH barycenters, the 371diffusion effect is avoided. Furthermore, new ensemble means, with respect to the GH 372distance, being placed side-by-side with deterministic forecasts provide useful information 373for forecasters in decision-making. 374

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A new view on the geometry of rainfall distributions should provide solutions for a 376 broader range of problems, not limited to ensemble means. We now try to tackle the 377reason underlying the resemblance of the ensemble means and the observations for 378 one-dimensional cumulative distributions in Fig. 3b that is left in the introduction. In the 379 metric space defined by the GH distance, GH barycenters are expected to approach 380 observations with increasing the number of ensemble members. Recall that the 3812-Wasserstein distance is equivalent to the Euclidean distance of inverses of cumulative 382distributions. This suggests that to grasp the convergence with respect to the GH distance, 383 a distance similar to the 2-Wasserstein distance, GH barycenters should be plotted in 384

cumulative forms. As expected, Fig. 7 shows that when the number of ensemble members increases from 20 (MEPS) to 100 (LETF100) and 1000 (LETKF1000) (see Duc et al., 2001, for detailed descriptions of the three ensemble prediction systems), the GH barycenters gradually approach the observations. However, what is more surprising is that the arithmetic means also converge to the GH barycenters. This explains why the arithmetic means converge to the observations in Fig. 3b.

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Since arithmetic means are, in fact, Euclidean barycenters, this raises a question on 392how we explain the convergence of two barycenters with increasing the number of 393 ensemble members. In general, it is easy to show a counter-example for this property, e.g., 394 many pairs of ensemble members with a fixed Wasserstein mean in Fig. 2b. Therefore, we 395hypothesize that this is a special property of rainfall ensemble forecasts in numerical 396weather prediction. In order to provide evidence for this hypothesis, Fig. 8 plots 397two-dimensional GH barycenters and arithmetic means from the same ensemble forecast 398systems in Fig. 7. Clearly, the arithmetic means again become nearly identical to the GH 399 barycenters when the number of ensemble members reaches 1000. Thus, 400two-dimensional rainfall distributions also show evidence for this property as in 401 one-dimensional distributions. Of course, this hypothesis should be verified for more 402 403 cases.

404

What are the other potential applications of the UOT-based geometry of rainfall 405 distributions? In this study, we use ensemble means to illustrate one of the potential 406 407 applications of the new geometry. However, it is important to verify the performance of GH barycenters in comparison with deterministic forecasts or traditional ensemble means. This 408 verification is usually quantified by objective verification scores as demonstrated in Fig. 6 409 with FSS. Due to its nature as a similarity measure, the GH distance should be a natural 410verification score in rainfall verification. Also, ensemble means do not make sense if 411 forecasts show a bi-modal probability distribution. In such cases, clustering needs to be 412deployed first, and clusters are represented by appropriate representatives. Then, the 413clustering can use the GH distance as a similarity measure, while clusters can be 414expressed by their corresponding GH barycenters. We will address these problems in the 415near future. 416

417

#### 418 Data Availability Statement

The datasets and source codes generated and/or analyzed in this study are available at
https://github.com/leducvn/uot4ens.

421

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428	
429	Appendix
430	A. Statistics of the forecast skill
431	In this section, all quantities are given in observation space. First, we suppose that an
432	ensemble forecast with K ensemble members is equivalent to K samples randomly

434 **Q** is the forecast error covariance. Thus, any ensemble member  $\mathbf{x}_k$  can be written under 435 the form

drawn from a multivariate normal distribution  $\mathcal{N}(\mathbf{x}_f, \mathbf{Q})$ , where  $\mathbf{x}_f$  is the forecast mean, and

436  $\mathbf{x}_k = \mathbf{x}_f + \mathbf{\varepsilon}_k$ , (A1)

437 where  $\mathbf{\epsilon}_k$  is a realization of forecast errors  $\mathcal{N}(\mathbf{0},\mathbf{Q})$ . The corresponding observation  $\mathbf{y}$  is 438 assumed to be the true state  $\mathbf{x}_t$ , which is unknown, contaminated by observation errors

439  $\mathbf{y} = \mathbf{x}_t + \mathbf{\varepsilon}_o,$  (A2)

440 where  $\mathbf{\epsilon}_o$  is a realization of observation errors  $\mathcal{N}(\mathbf{0},\mathbf{R})$ , and  $\mathbf{R}$  is the observation error 441 covariance.

442

433

We define the forecast skill to be the difference between the ensemble mean  $\bar{\mathbf{x}} = \sum \mathbf{x}_k / K$ and the observation  $\mathbf{y}$ 

445 
$$\mathbf{s} = \overline{\mathbf{x}} - \mathbf{y} = \mathbf{x}_f - \mathbf{x}_t + \frac{1}{K} \Sigma \boldsymbol{\varepsilon}_k - \boldsymbol{\varepsilon}_o,$$
 (A3)

where we use (A1) and (A2) to obtain the last expression. Clearly, s is a random variable,

- 447 and it is straightforward to calculate its statistics
- 448  $\mathbf{E}[\mathbf{s}] = \mathbf{x}_f \mathbf{E}[\mathbf{x}_t]. \tag{A4}$
- 449  $\operatorname{E}[\mathbf{s}\mathbf{s}^{\mathrm{T}}] = \operatorname{E}[(\mathbf{x}_{f} \mathbf{x}_{t})(\mathbf{x}_{f} \mathbf{x}_{t})^{\mathrm{T}}] + \frac{1}{K}\mathbf{Q} + \mathbf{R}.$  (A5)

450 The right-hand sides of (A4) and (A5) are derived using the fact that  $\mathbf{x}_t, \boldsymbol{\varepsilon}_k, \boldsymbol{\varepsilon}_o$  are 451 independent.

452

In order to get the final forms of (A4) and (A5), we need to make an assumption on the true state  $\mathbf{x}_t$ . Since  $\mathbf{x}_f$  is the mode of the probability distribution of the forecasts  $\mathcal{N}(\mathbf{x}_f, \mathbf{Q})$ , we assume that  $\mathbf{x}_f$  is a good approximation of  $\mathbf{x}_t$ 

456  $\mathbf{x}_t \approx \mathbf{x}_f$ . (A6)

457 This means rather than a random variable,  $\mathbf{x}_t$  is considered as a fixed quantity given

approximately by  $\mathbf{x}_{f}$ . Under this assumption, (A4) and (A5) respectively become

 $459 \quad \mathbf{E}[\mathbf{s}] = \mathbf{0}. \tag{A7}$ 

460  $\operatorname{E}[\mathbf{s}\mathbf{s}^{\mathrm{T}}] = \frac{1}{K}\mathbf{Q} + \mathbf{R}.$  (A8)

461 If observation errors **R** are negligible compared to forecast errors **Q**, (A8) points out that  $\bar{\mathbf{x}}$ 462 asymptotically converges to  $\mathbf{y}$  with the error being proportional to  $1/\sqrt{K}$ .

463

464 The assumption (A6) can be further relaxed only by requiring  $\mathbf{x}_t$  to follow the same

465 probability distribution of the forecasts 
$$\mathcal{N}(\mathbf{x}_{f}, \mathbf{Q})$$

466 
$$\mathbf{x}_t = \mathbf{x}_f + \boldsymbol{\varepsilon}_t.$$
 (A9)

where  $\varepsilon_t$  is a realization of forecast errors  $\mathcal{N}(\mathbf{0}, \mathbf{Q})$ . This means  $\mathbf{x}_t$  is considered as a random variable now, and is indistinguishable from all ensemble members. Under this assumption, (A4) and (A5) respectively become

- 470 E[s] = 0. (A10)
- 471  $\operatorname{E}[\mathbf{s}\mathbf{s}^{\mathrm{T}}] = \frac{K+1}{K}\mathbf{Q} + \mathbf{R}.$  (A11)

472 Again, if observation errors **R** are negligible, (A11) represents the well-known spread-skill

473 relationship in ensemble forecast.

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